

Mean Energy Density of Photogenerated Magnetic Fields Throughout the EoR

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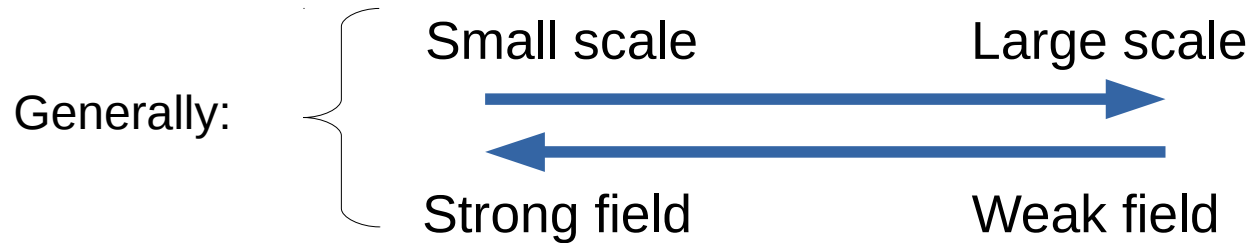
Durrive & Langer, 2015, MNRAS, arXiv:1506.08177

➔ Durrive, Tashiro, Langer, Sugiyama 2017, MNRAS, 472, 1649

Durrive & Aubert, 2017, in prep

Peering towards Cosmic Dawn, IAU Symposium 333, 2-6/10/17, Dubrovnik

- Magnetic fields **everywhere**: from stars to galaxies to cosmic voids



- **Cosmological** Magnetic fields:

Observational upper bounds:

CMB: $B < 5$ nG (comoving) at 1 Mpc ([Planck results 2015: XIX](#))

Constraints from structure formation $B < \text{nG}$ at protogalactic scales
([Wasserman 1978](#), [Kim et al 1996](#))

Observational **lower** bounds:

High energy gamma rays (Fermi and HESS): $B > 10^{-16}$ or 10^{-18} G (?)
in a significant fraction of the IGM
([Neronov&Vovk 2010](#), [Taylor et al 2011](#), [Takahashi et al 2011](#), ...)

Origin(s)?

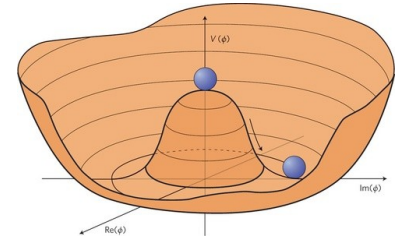
- Current paradigm:
 - 1) Generate **weak seeds**
 - 2) **Amplified**: compression during structure formation (flux freezing) + dynamos
- Turbulence in structures \rightarrow B fields lost their initial properties
 \rightarrow look at the **Intergalactic medium** where seeds did not evolve too much

- Current status of amplification process studies:

→ we need $\sim 10^{-22}$ to 10^{-12} G seeds

- Numerous mechanisms:

(Reviews see e.g. Ryu et al 2012, Widrow et al 2012, Durrer & Neronov 2013, ...)

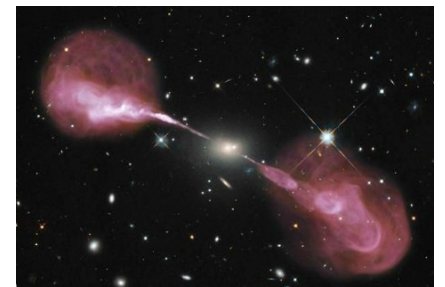


I) Primordial Universe mechanisms

- Inflation: quantum fluctuations of electromagnetic field, but need non-standard electromagnetism
- Phase transitions: electroweak and quark-hadron
- Recombination: rotating plasma blobs interacting with background radiation

II) Post recombination mechanisms

- Thermal (Biermann) battery: in stars, from cosmological shocks during cosmic web formation, from propagating ionization fronts at EoR in large structures
- Plasma instabilities: many, but e.g. Weibel instability
- Radiation: Thomson scattering In protogalaxies, **Photoionization** at EoR in the IGM or around first stars
- Outflows: Galactic winds from galaxies in clusters, from void galaxies, AGN outflows

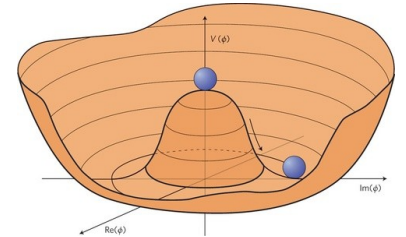


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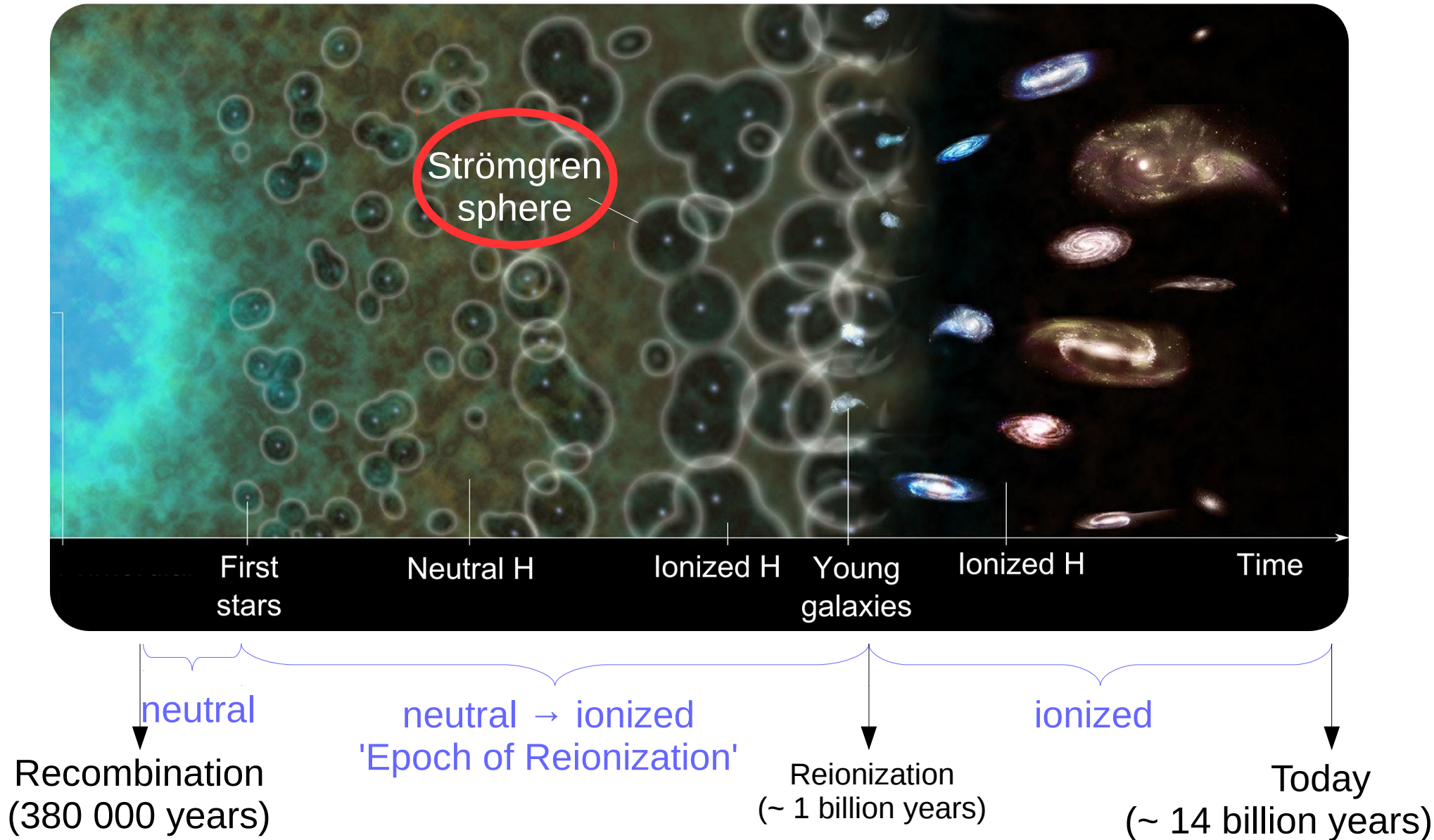
II) Post recombination

**No preferred mechanism so far.
Fields not strong enough on intergalactic scales**

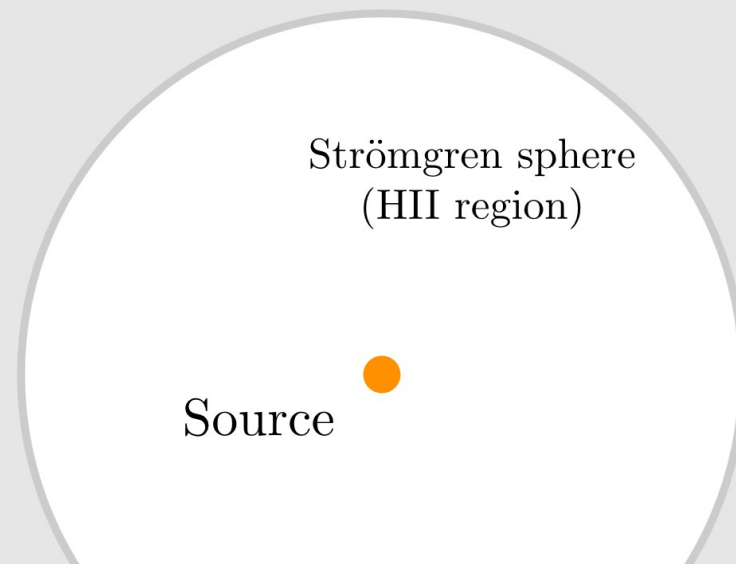


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Astrophysical mechanism generating **intergalactic** magnetic fields at the Epoch of Reionization



Intergalactic medium (Hydrogen)



Strömgren sphere
(HII region)

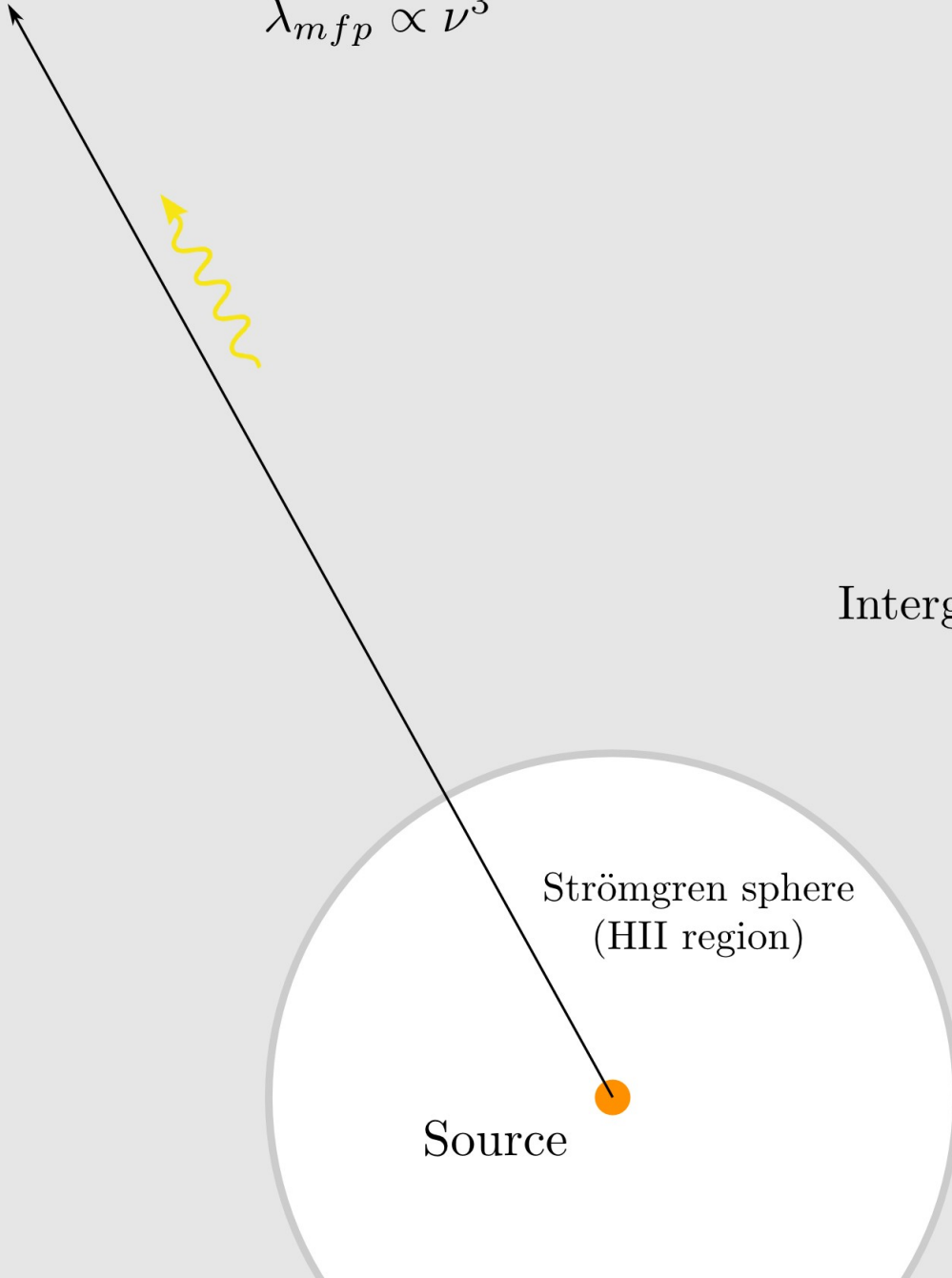
Source

$$\lambda_{mfp} \propto \nu^3$$

Intergalactic medium (Hydrogen)

Strömgren sphere
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Source



$$\lambda_{mfp} \propto \nu^3$$

Charge separation by photoionization

Intergalactic medium (Hydrogen)

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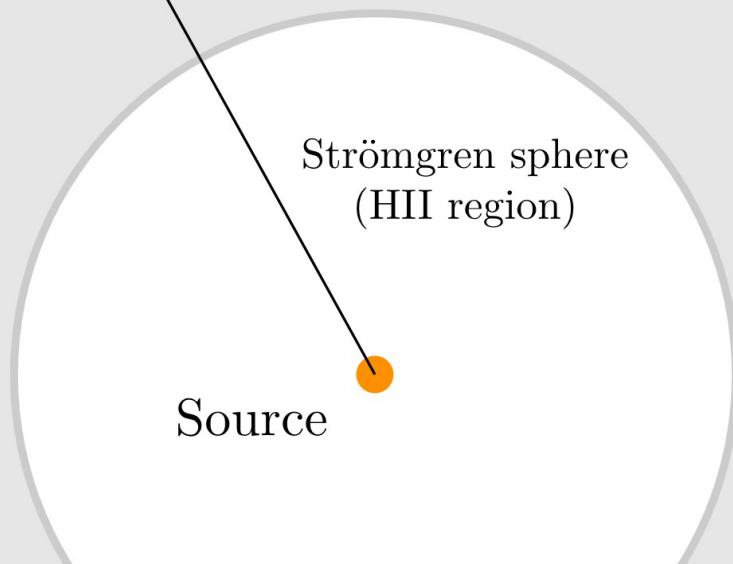
Source



The diagram illustrates the Strömgren sphere, a region of ionized hydrogen (HII) surrounding a source. A black line with an arrow at its tip represents the ionization front, extending from the source (an orange dot) into the intergalactic medium (Hydrogen). Along this front, a blue dot and a yellow wavy arrow indicate charge separation by photoionization. The equation $\lambda_{mfp} \propto \nu^3$ is shown above the front. The region inside the sphere is labeled 'Strömgren sphere (HII region)' and the region outside is labeled 'Intergalactic medium (Hydrogen)'. The source is labeled 'Source'.

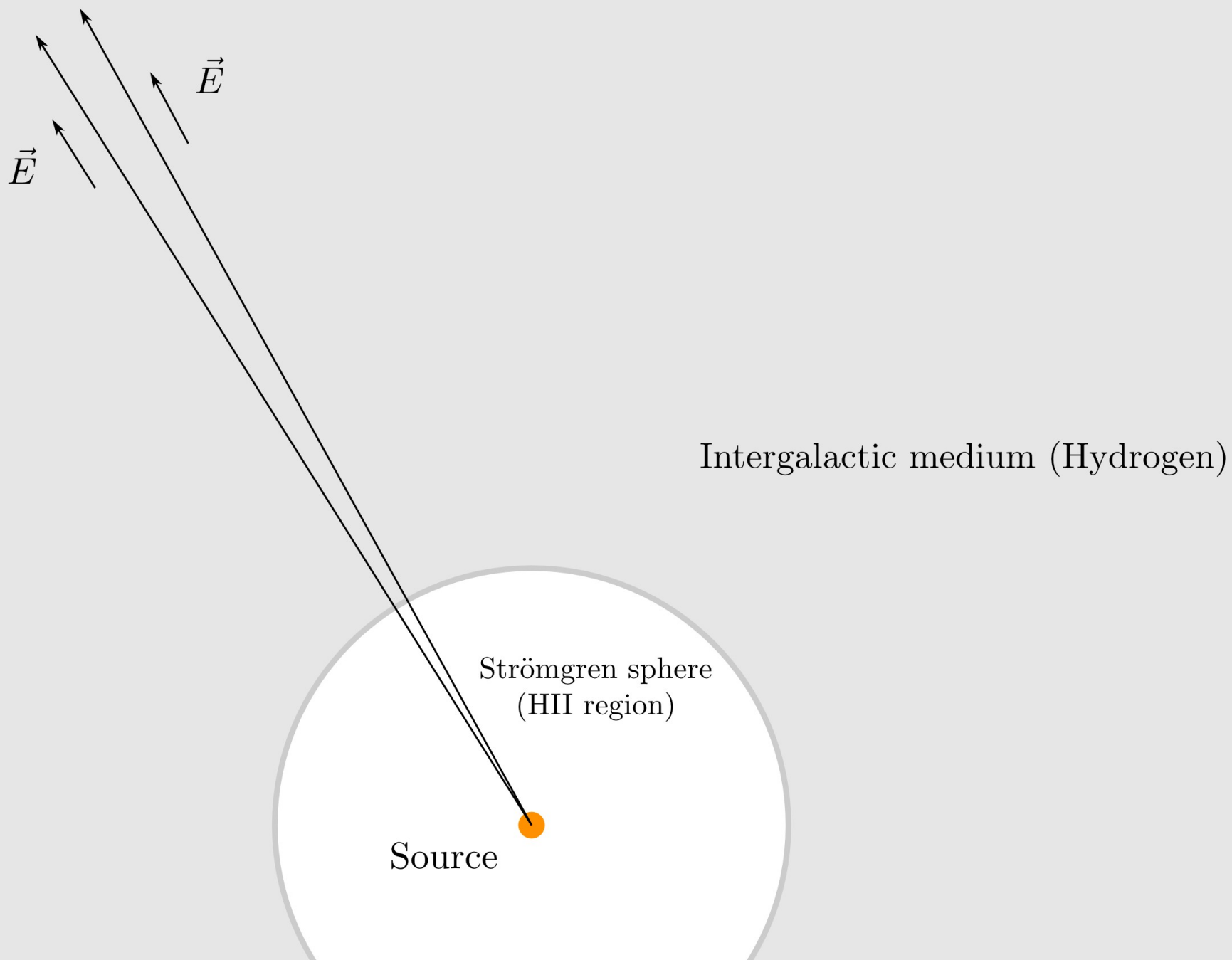


Intergalactic medium (Hydrogen)

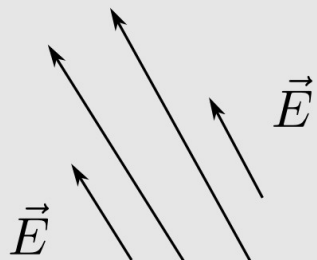


Strömgren sphere
(HII region)

Source



$$\vec{\nabla} \times \vec{E} = \vec{0}$$



Intergalactic medium (Hydrogen)

Strömgren sphere
(HII region)

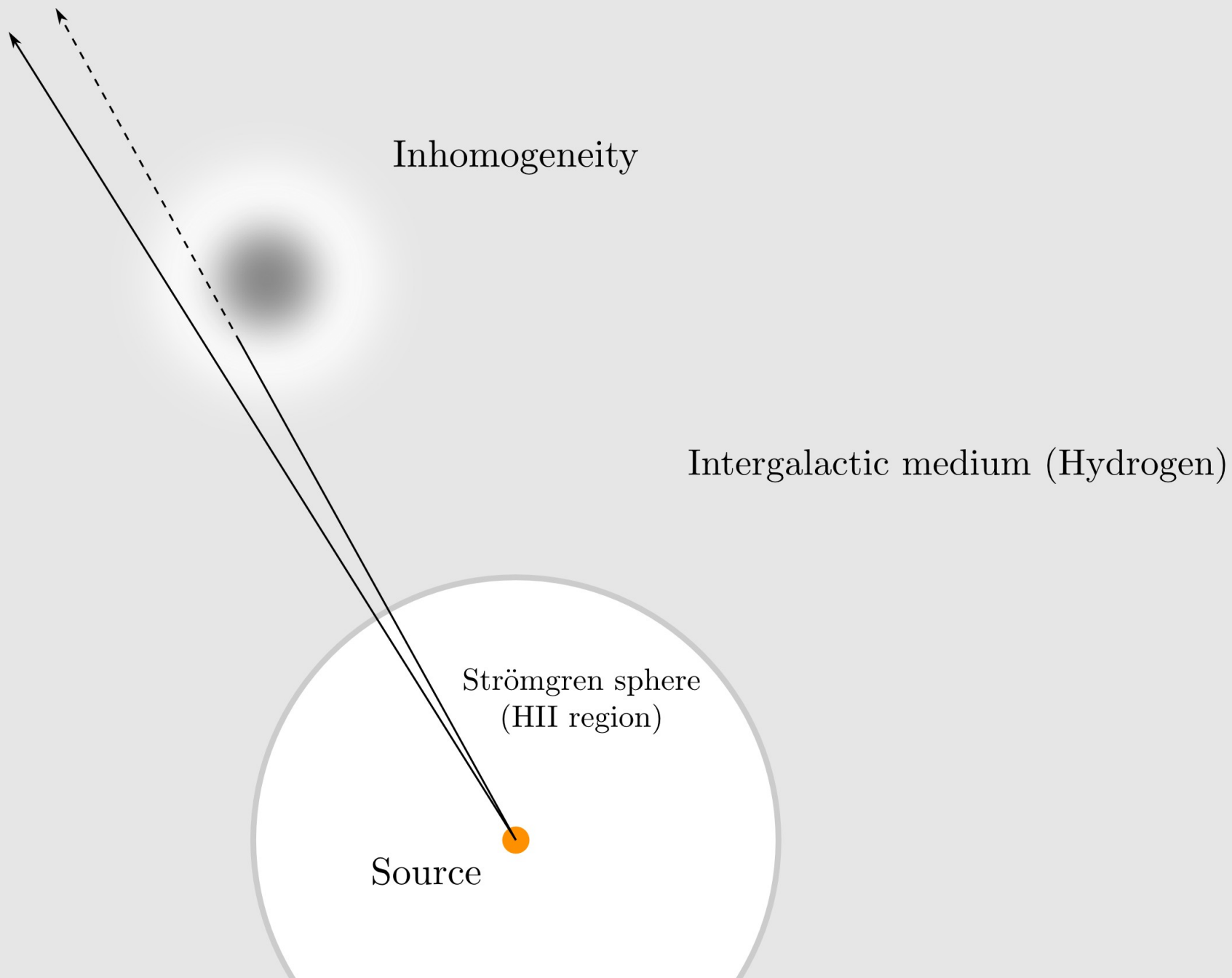
Source

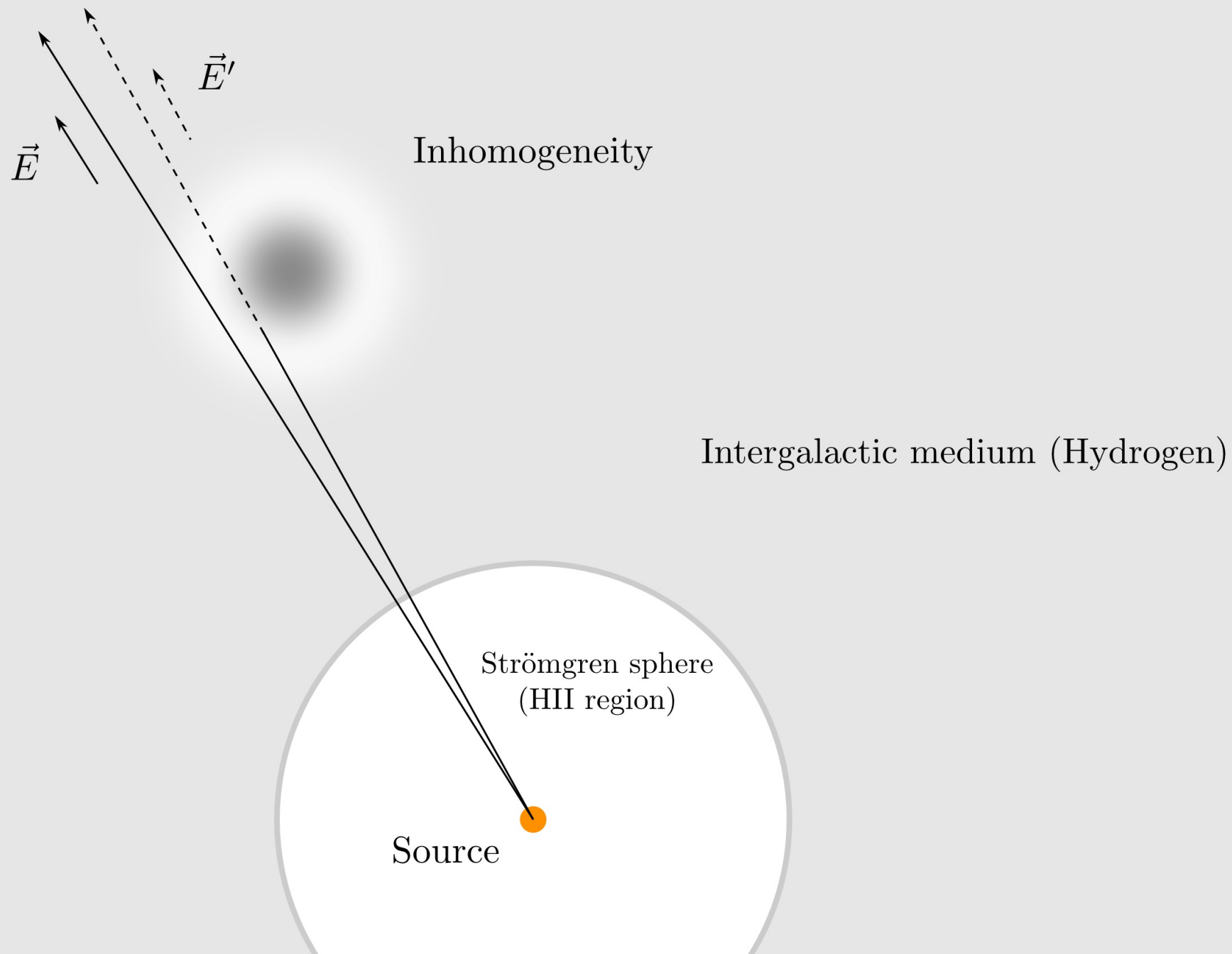
Maxwell-Faraday equation: $\partial_t \vec{B} = -c \vec{\nabla} \times \vec{E}$

→ Need **rotational E field** to generate B

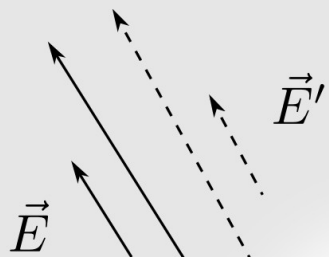
Inhomogeneities of Intergalactic medium enable this

(Langer et al 2005)



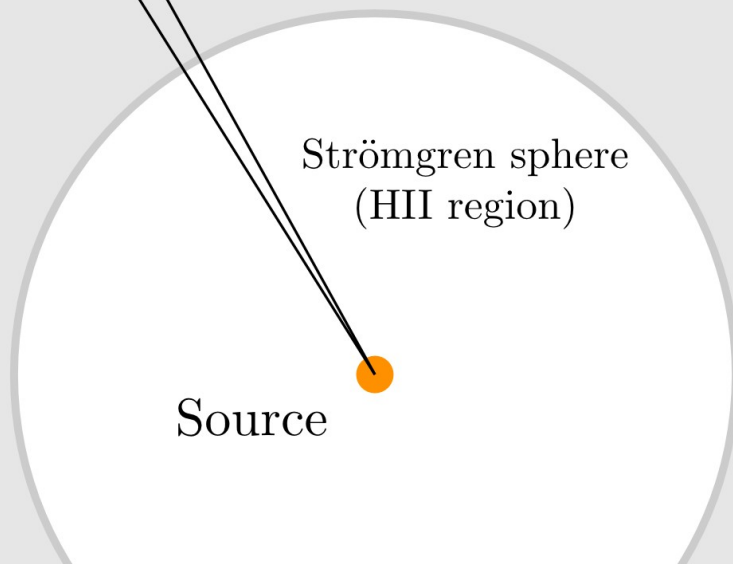


$$\vec{\nabla} \times \vec{E} \neq \vec{0}$$



Inhomogeneity

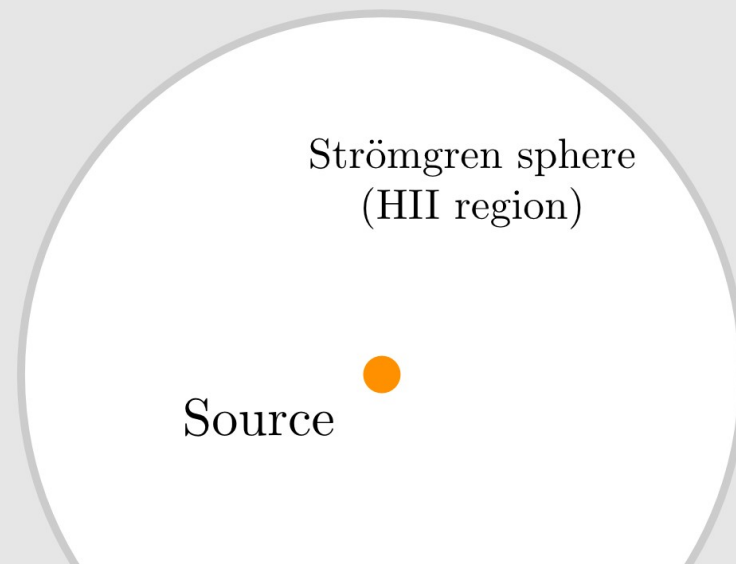
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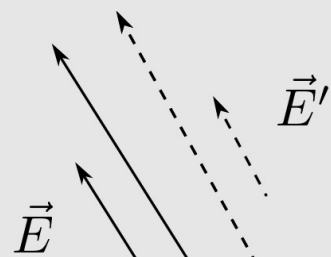
Intergalactic medium (Hydrogen)

Anisotropic HII region

Source



$$\vec{\nabla} \times \vec{E} \neq \vec{0}$$



Intergalactic medium (Hydrogen)

Anisotropic HII region

Source



Formally

(Durrive & Langer, 2015, MNRAS)

Photoionization = local modification of the number of electrons
and of their **velocity distribution**

⇒ Kinetic theory!
Source term in Boltzmann equation
of electron distribution function:

$$\frac{df}{dt} = \partial_t f|_{\text{photoionization}}$$

Momentum transferred
from photons to electrons:

$$m_e \vec{v} = f_{mt}(\nu) \frac{h\nu}{c} \hat{r}$$

Fraction of momentum transferred

Momentum transfer rate

Induction equation:

$$\partial_t \vec{B} = -\frac{c}{e} \frac{\vec{\nabla} n_e \times \vec{\nabla} p_e}{n_e^2} - \frac{c}{e} \vec{\nabla} \times \frac{\dot{\vec{p}}_e}{n_e}$$

Biermann

**Photoionization:
Radiation induces magnetic fields!**

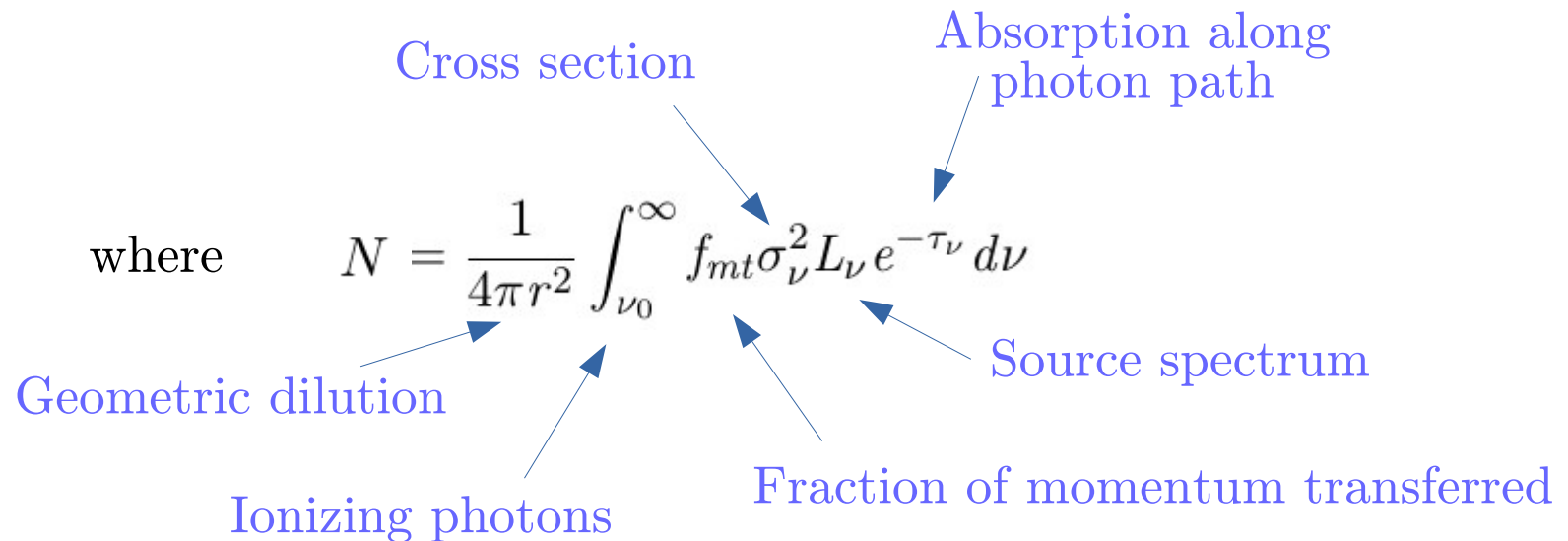
Resulting magnetic field

(Durrive & Langer, 2015, MNRAS)

Source of B :
Anisotropy of the column density

$$\vec{B}(t, \vec{r}) = t \frac{N}{ex_e} \vec{\nabla} \int_{r_s}^r n_{HI} dr \times \hat{r}$$

where

$$N = \frac{1}{4\pi r^2} \int_{\nu_0}^{\infty} f_{mt} \sigma_{\nu}^2 L_{\nu} e^{-\tau_{\nu}} d\nu$$


Cross section

Absorption along photon path

Source spectrum

Fraction of momentum transferred

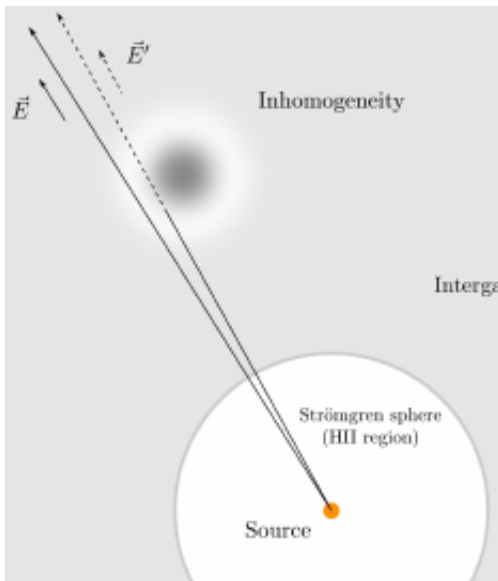
Ionizing photons

Geometric dilution

Resulting magnetic field

(Durrive & Langer, 2015, MNRAS)

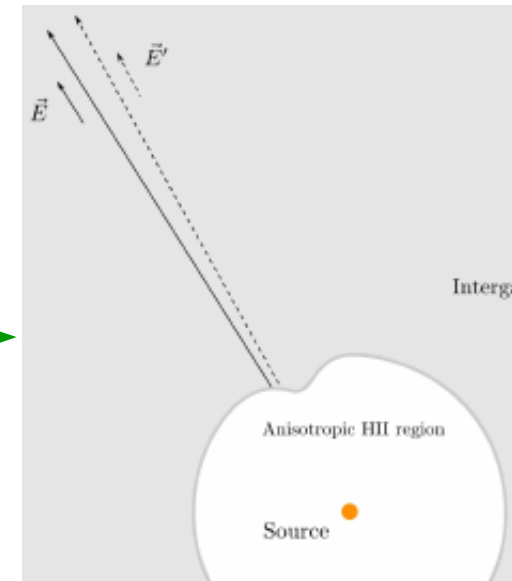
Source of B :
Anisotropy of the column density



Inhomogeneity

$$\vec{B}(t, \vec{r}) = t \frac{N}{ex_e} \vec{\nabla} \int_{r_s}^r n_{HI} dr \times \hat{r}$$

Strömgren radius



Previously: Study for various isolated sources at various epochs:

PopIII clusters: $B \sim 10^{-21} \text{ G}$ to 10^{-18} G on $\sim 2 \text{ kpc}$

First galaxies: $B \sim 10^{-21} \text{ G}$ to 10^{-19} G on $\sim 20 \text{ kpc}$

Quasars: $B \sim 10^{-23} \text{ G}$ to 10^{-21} G on $\sim 2 \text{ Mpc}$

New work: In the cosmological context of EoR:

Analytic approach:

Durrive, Tashiro, Langer, Sugiyama, 2017

Numerical approach:

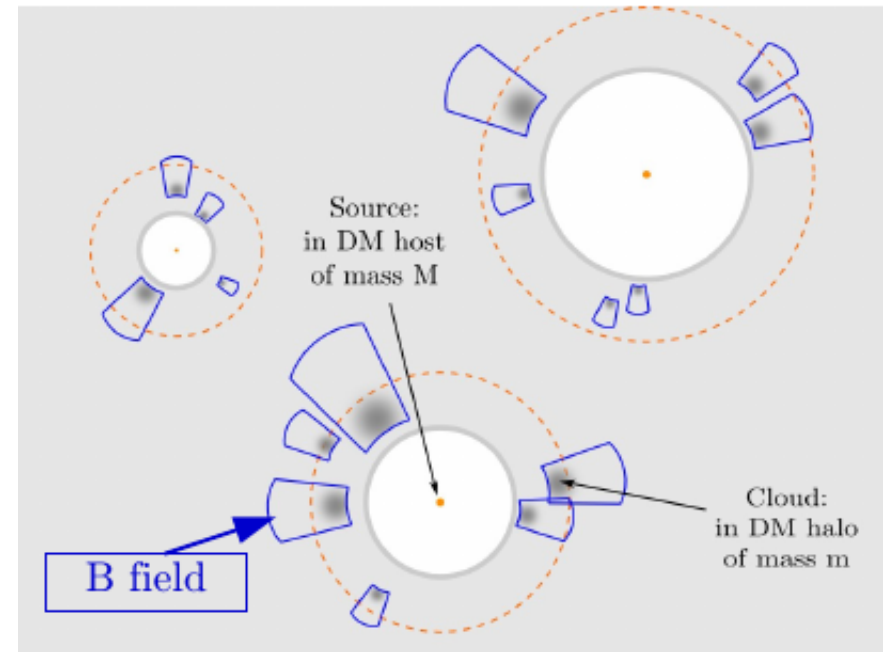
Durrive, Aubert, in prep

Global magnetization level of the Universe

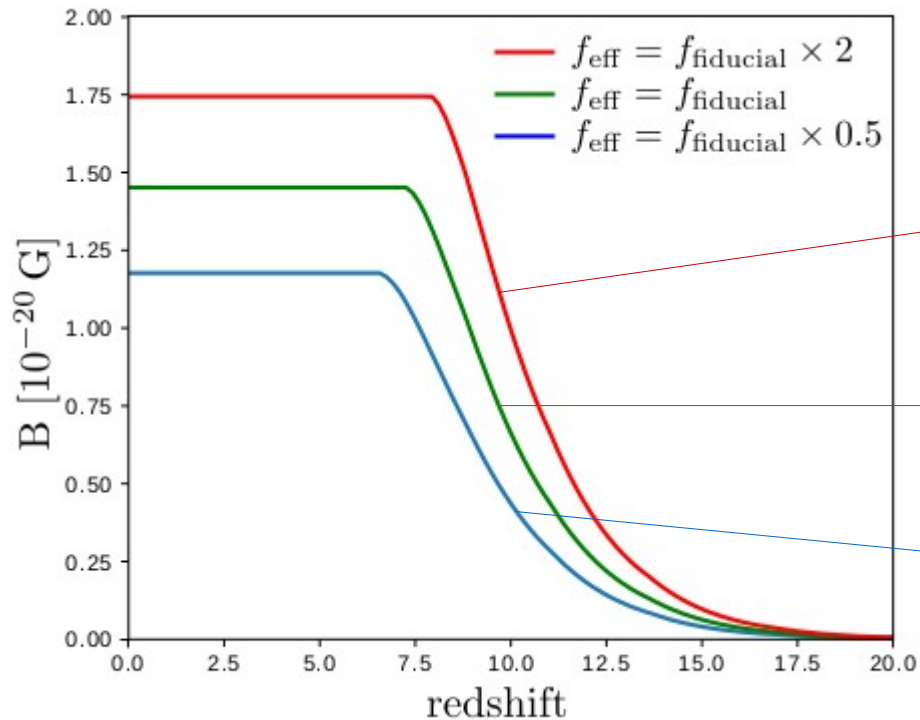
(Durrive, Tashiro, Langer, Sugiyama 2017, MNRAS)

Case of galaxies:

Distribution of sources & clouds:
Given by underlying **Dark Matter halos**
(Press-Schechter formalism)



Mean magnetic field in the Universe:



Universe with '**strongly ionizing**' galaxies
(**maximal** escape fraction & stars formed
to stay consistent with Planck)

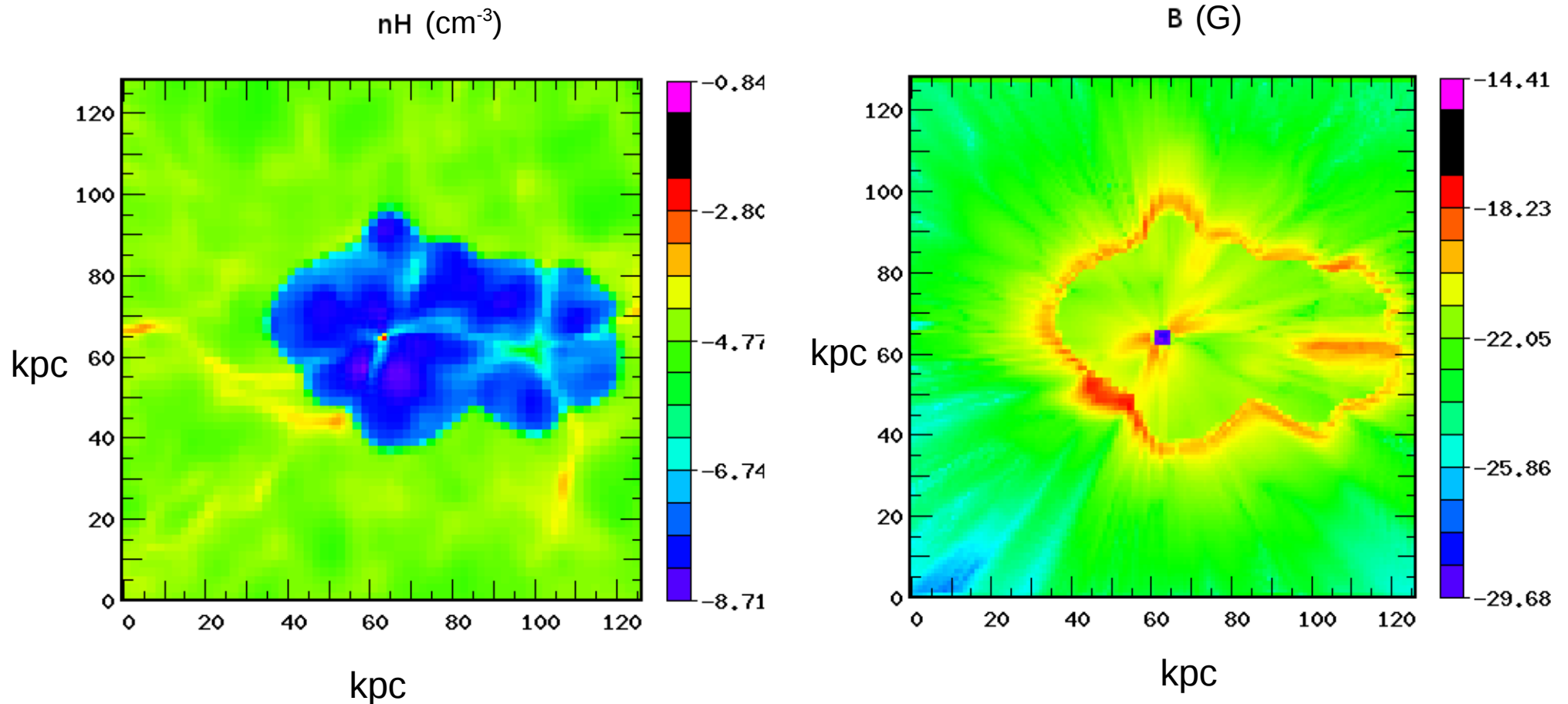
Fiducial model

Universe with '**weakly ionizing**' galaxies
(**minimal** escape fraction & stars formed
to stay consistent with Planck)

Numerical approach

(Durrive & Aubert, 2017, in prep)

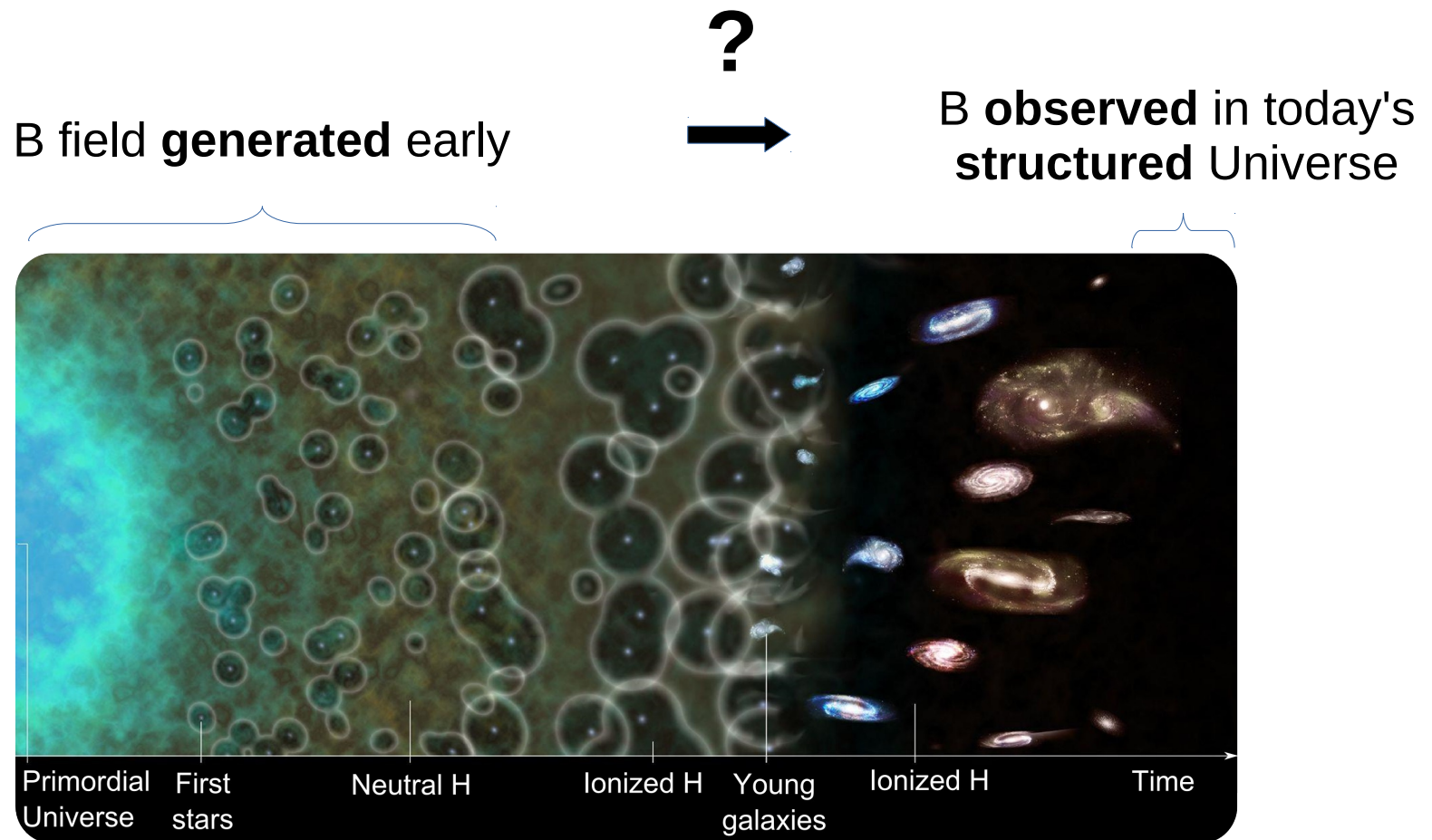
Generated B field with realistic profiles from cosmological simulations:



(example of a primordial galaxy at $z = 10$)

Ongoing work!

Evolution in the cosmic web?



→ Need to study the **evolution** of cosmological magnetic fields

Conclusion and discussion

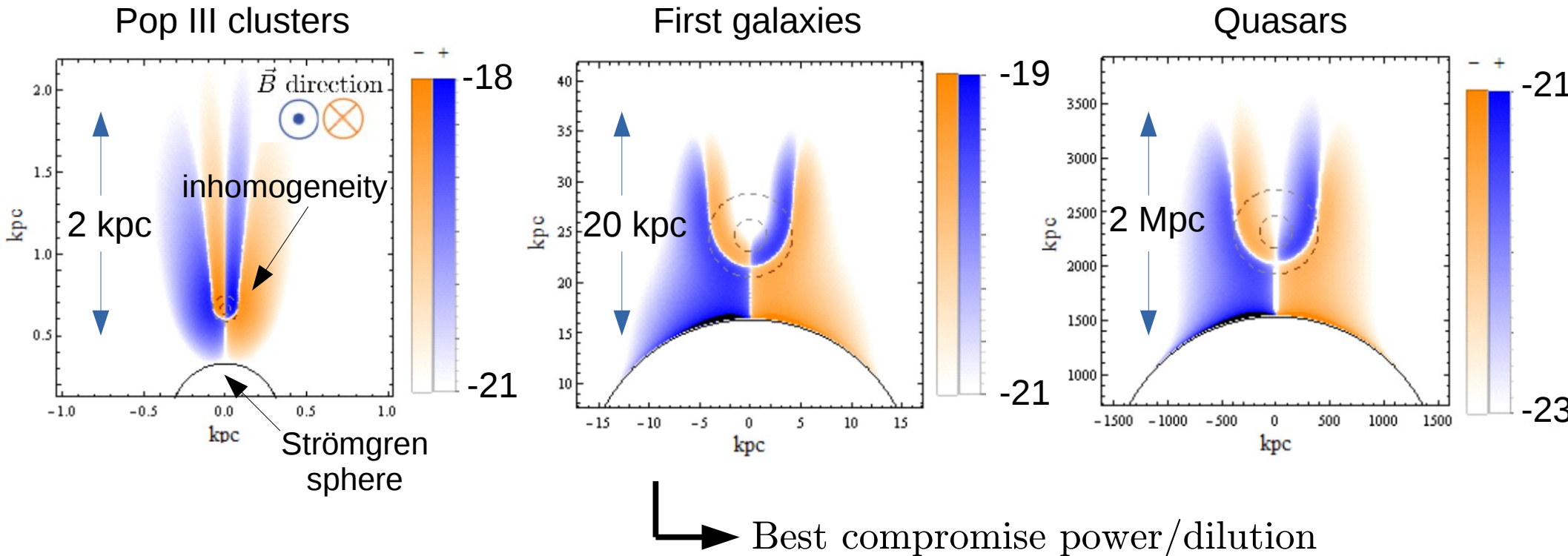
- Astrophysical mechanism, operating for any source, **all along the EoR**
- Strengths comparable to Biermann battery, but on entire inter-source scales
 - ⇒ Contributes to **magnetization of the whole Intergalactic medium**
interesting for voids!
- Specific spatial configuration:
 - may help discriminate the seeds from other mechanisms
- Directly measurable seeds ?
 - 10^{-19} G fields prior and during EoR
(Venumadhav et al 2017, Gluscevic et al 2017)
 - SKA
- Evolution of cosmological B fields in the cosmic web?
(e.g. Vazza et al 2014)

Thank you for your attention

Typical spatial distributions and scales

(Durrive & Langer, 2015, MNRAS)

- Gaussian inhomogeneities \rightarrow analytical expressions
- Explicit lengthscales & magnetized regions
- Studied properties for various sources at various epochs:



- Compared with intersource distances:
 \Rightarrow magnetization of the **whole intergalactic medium**

