

# The Neutral Islands during the Late Epoch of Reionization

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# Outline

- \* Introduction:
  - \* Cosmic reionization
  - \* The excursion set theory of reionization
- \* The Island Model
  - \* The bubbles-in-island effect
  - \* The ionizing background
- \* The islandFAST - semi-numerical simulation of late EoR
- \* Results
  - \* Size distribution & evolution of neutral islands
  - \* The evolution of ionizing background
- \* Summary and Outlook

# What is the Reionization Era?

## A Schematic Outline of the Cosmic History

Time since the Big Bang (years)

~ 300 thousand

~ 500 million

~ 1 billion

~ 9 billion

~ 13 billion



← The Big Bang

The Universe filled with ionized gas

← The Universe becomes neutral and opaque

The Dark Ages start

Galaxies and Quasars begin to form  
The Reionization starts

The Cosmic Renaissance  
The Dark Ages end

← Reionization complete, the Universe becomes transparent again

Galaxies evolve

The Solar System forms

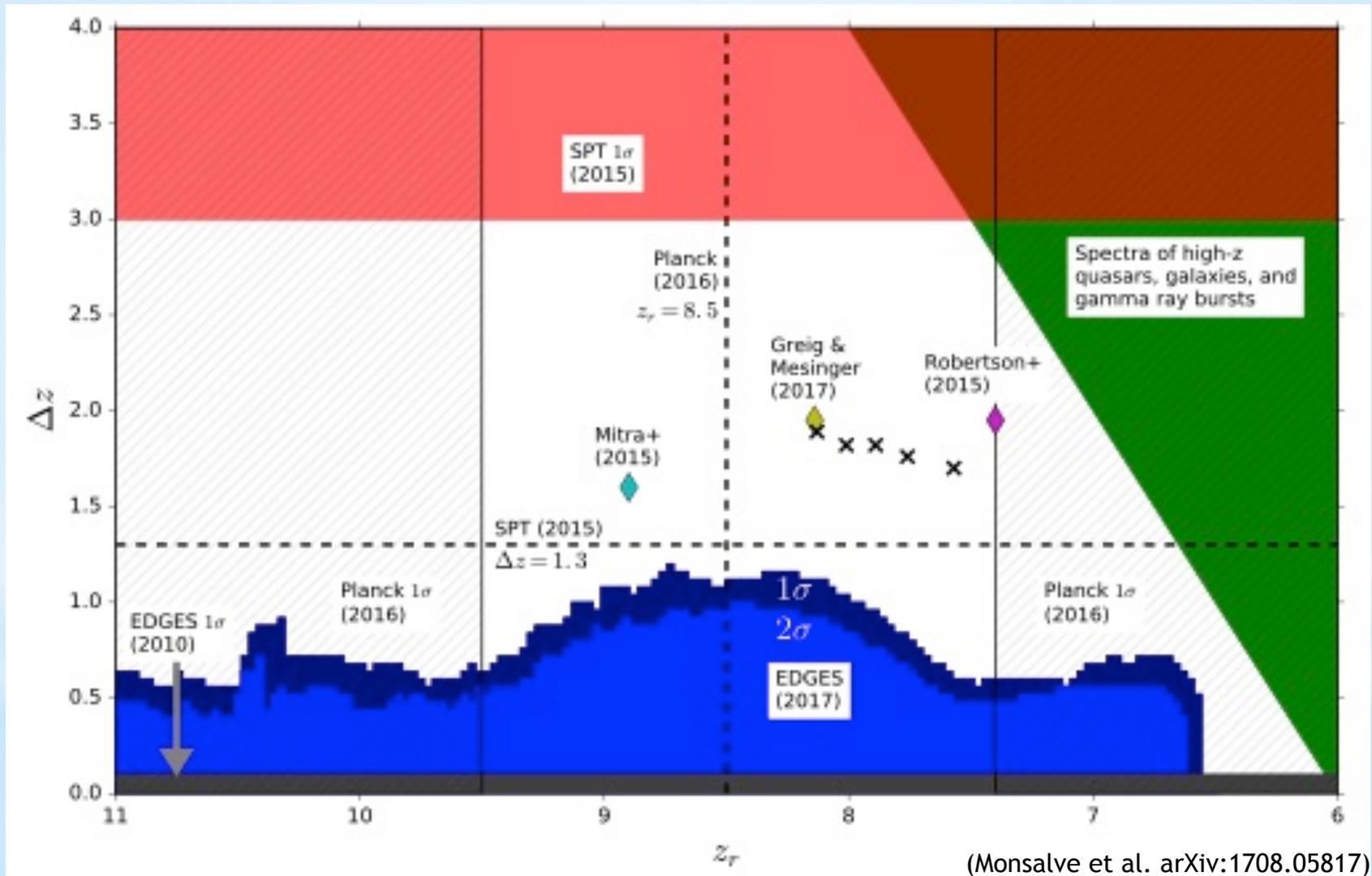
Today; Astronomers figure it all out!

Neutral IGM after recombination

Epoch of reionization (EoR)

Ionized IGM indicated by quasar absorption spectra

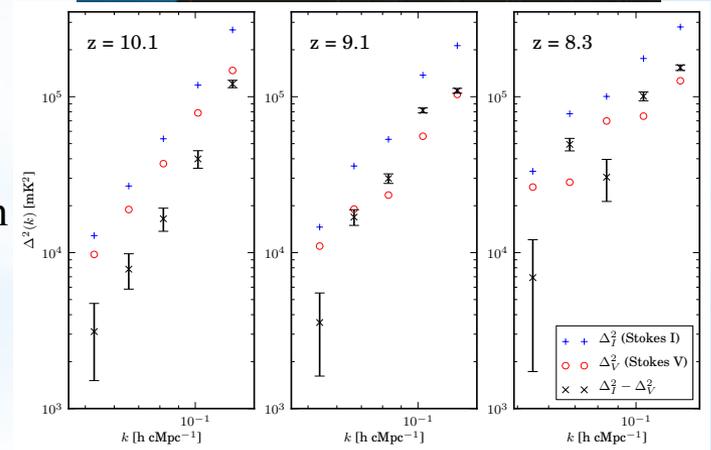
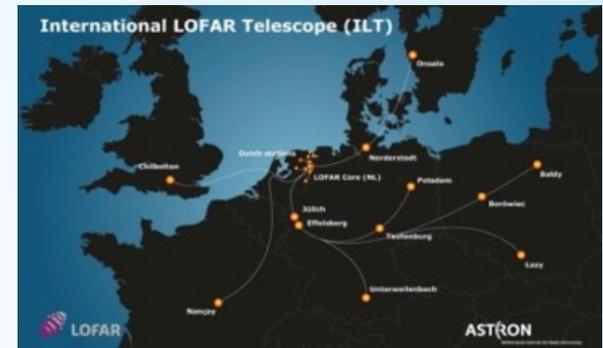
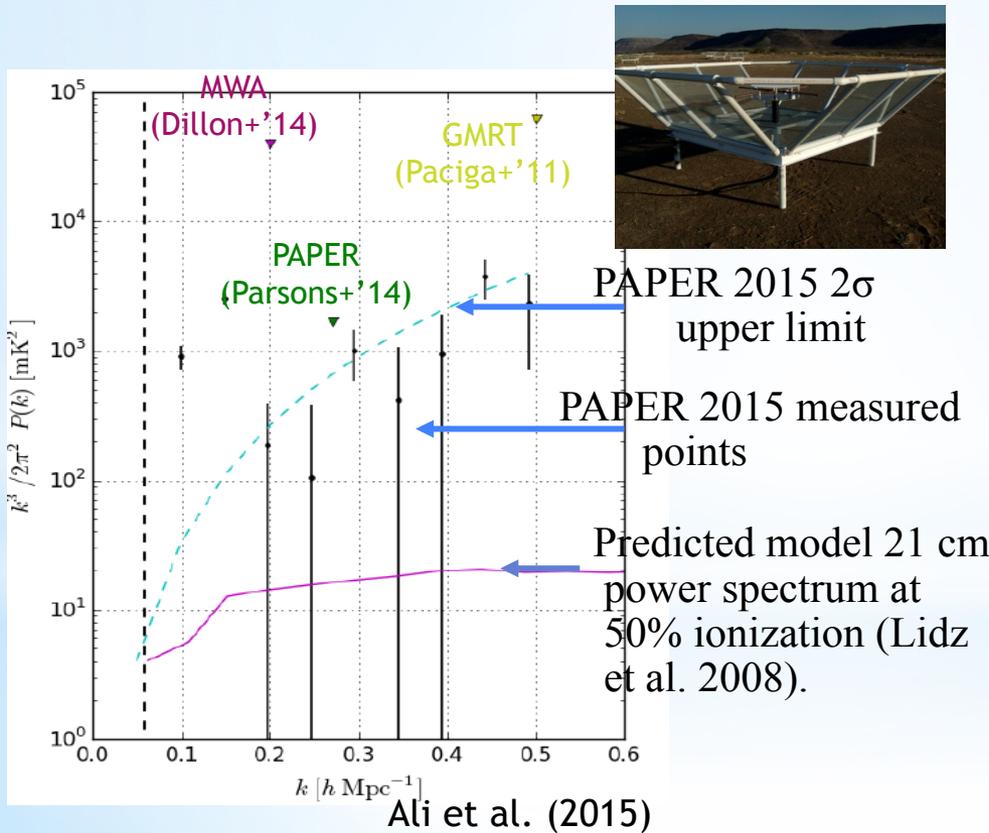
# Observational Constraints



(Monsalve et al. arXiv:1708.05817)

# Observational Constraints

- \* Upper limit on the 21cm power spectrum:
  - \* PAPER →  $(22.4 \text{ mK})^2$  at  $k=0.15 - 0.5 \text{ h Mpc}^{-1}$  at  $z=8.4$
  - \* LOFAR →  $(79.6 \text{ mK})^2$  at  $k = 0.053 \text{ h cMpc}^{-1}$  in the range  $z = 9.6 - 10.6$ .



Patil et al. (2017)

# The upcoming 21 cm experiments



SKA-low



Hydrogen Epoch of Reionization Array  
(HERA)



# Analytical models of reionization

\* Early stage – the “bubble model” (Furlanetto et al. 2004)

-- growing ionized bubbles

\* Late stage – the “island model” (Xu et al. 2014)

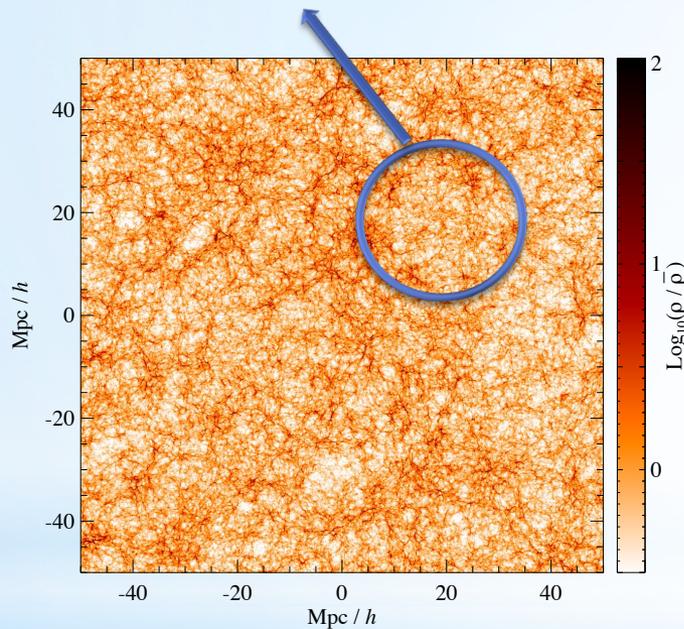
-- shrinking neutral islands

**→ Both based on the excursion set theory**

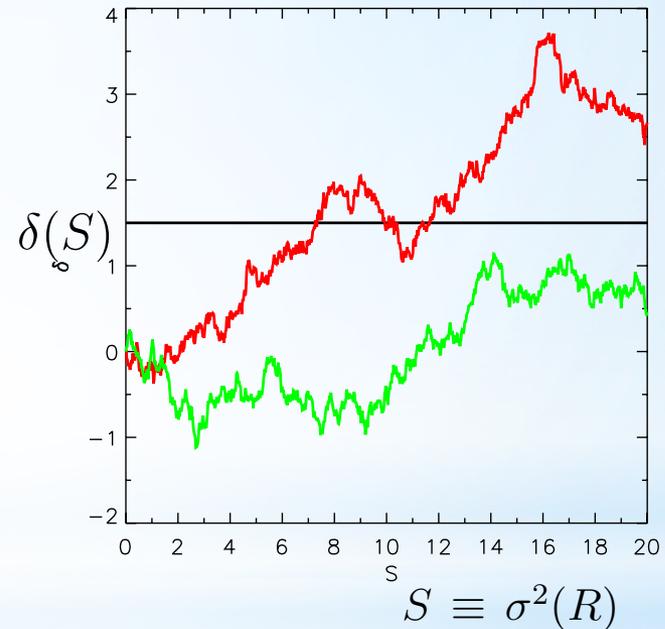
# The Excursion Set Theory of Halo Model

(Bond et al. 1991, Lacey & Cole 1993)

$$\delta(\vec{x}; R) \equiv [\rho(\vec{x}) - \rho_M] / \rho_M$$



- Halo density barrier

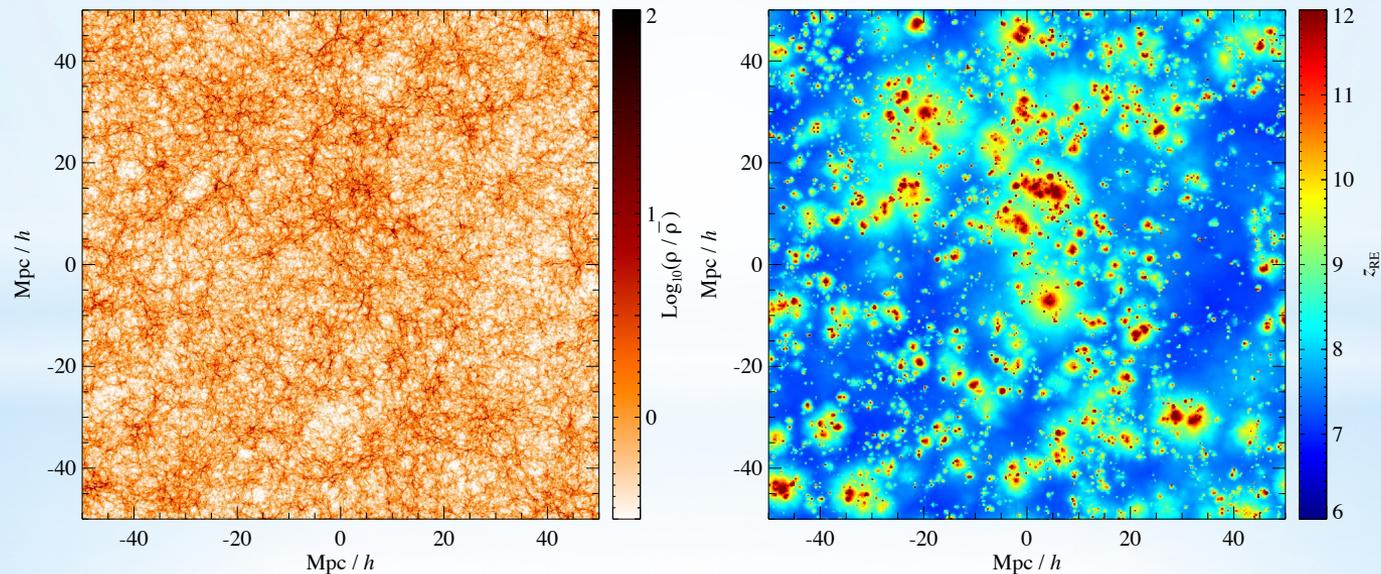


- Solving a diffusion equation  $\rightarrow$  “*first-crossing distribution*”  
 $\rightarrow$  halo mass function

# Why excursion set theory?

→ Full RT-simulations are computationally expensive

→ The reionization field follows the density field on large scales (Battaglia et al. 2013)



(From Battaglia et al. 2013 ApJ, 776, 81)

# The Excursion Set Approach for ionized bubbles

## - The bubble model of reionization

(Furlanetto et al. 2004)

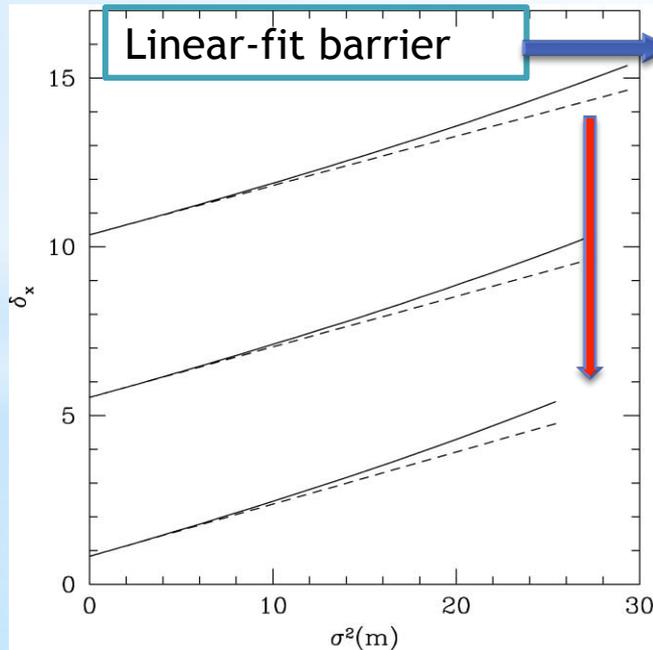
- \* Relate the ionization field to the initial density field
- \* Ask whether an isolated region of mass  $M$  can be fully self-ionized.

$$f_{\text{coll}} \geq f_x \equiv \zeta^{-1}.$$

$$\delta_m \geq \delta_x(m, z) \equiv \delta_c(z) - \sqrt{2}K(\zeta)[\sigma_{\text{min}}^2 - \sigma^2(m)]^{1/2}$$

$$K(\zeta) = \text{erf}^{-1}(1 - \zeta^{-1}).$$

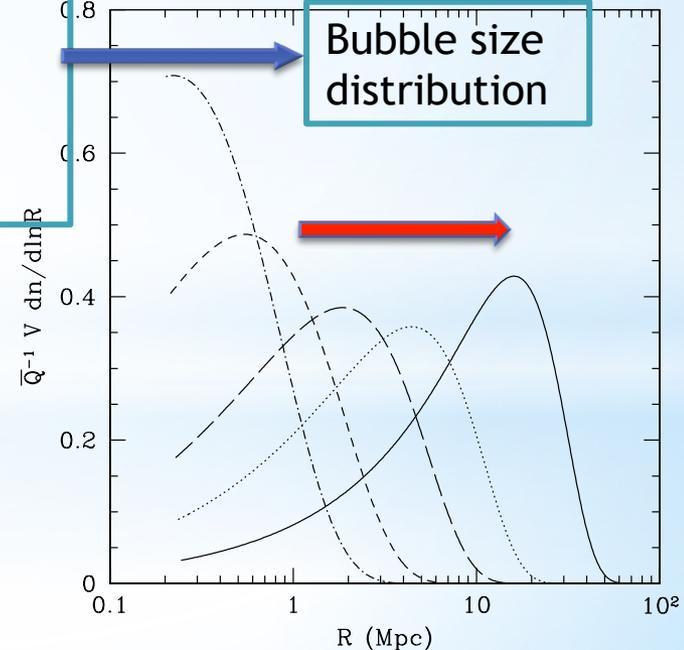
The bubble barrier



Linear-fit barrier

First-up-crossing distribution (analytical)

*Bubble-in-bubble problem*



Bubble size distribution

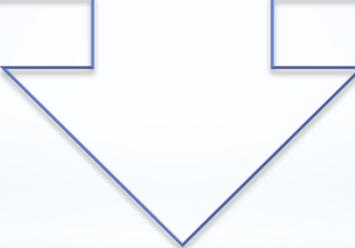
## However, after percolation...

1. The isolated and spherical assumption for the ionized bubbles breaks down

→ the neutral islands are more isolated

2. The existence of an ionizing background

→ the shape of barriers could be changed  
(the linear fit may not apply)



## The island model

It would be relatively easier for the upcoming instruments to probe the signal at the late reionization stages.

# The Island Model

(Xu et al. 2014)

- \* Negative island barrier (“inside-out” reionization)
- \* Island mass scales are identified by *first-down-crossings* through the island barrier (but not the “never-up-crossing” distribution).
- \* With the inclusion of an ionizing background, the condition of keeping from being ionized:

$$\xi f_{\text{coll}}(\delta_M; M, z) + \frac{\Omega_m}{\Omega_b} \frac{N_{\text{back}} m_{\text{H}}}{M X_{\text{H}} (1 + \bar{n}_{\text{rec}})} < 1,$$

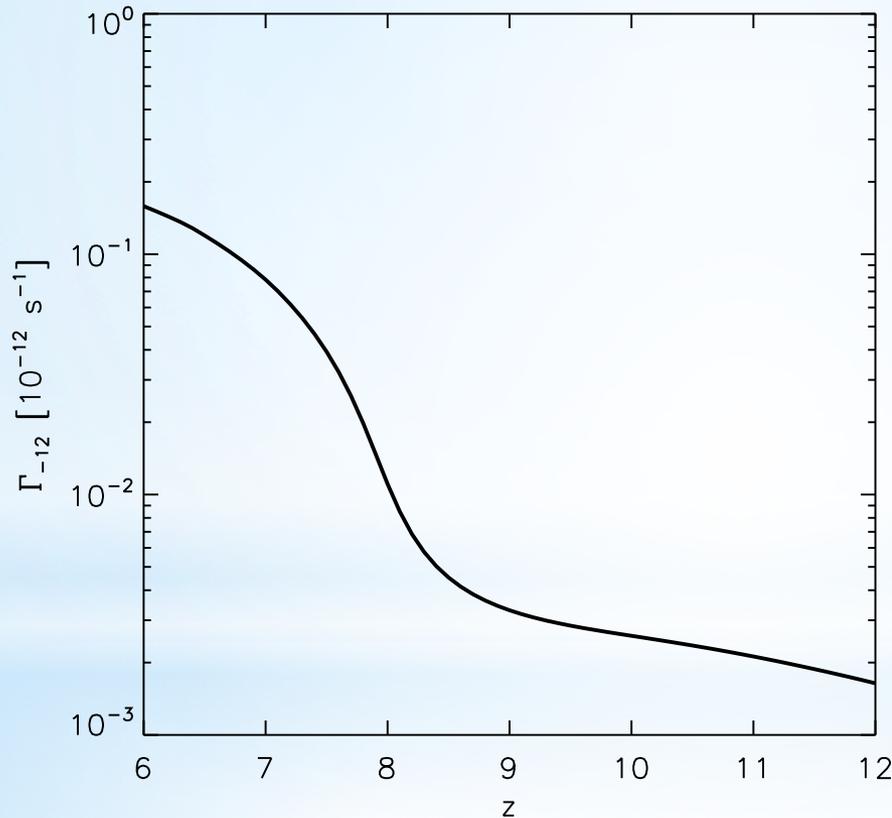
→ The island barrier:

The contribution of background ionizing photons

$$\delta_M < \delta_{\text{I}}(M, z) \equiv \delta_c(z) - \sqrt{2[S_{\text{max}} - S(M)]} \operatorname{erfc}^{-1} [K(M, z)],$$

$$K(M, z) = \xi^{-1} \left[ 1 - N_{\text{back}} (1 + \bar{n}_{\text{rec}})^{-1} \frac{m_{\text{H}}}{M (\Omega_b / \Omega_m) X_{\text{H}}} \right].$$

# The ionizing background



- \* Considering the effect of *Lyman limit systems* on the mean free path of ionizing photons
- \* Scaling the hydrogen photoionization rate to be  $\Gamma_{\text{HI}} = 10^{-12.8} \text{ s}^{-1}$  at redshift 6, as suggested by recent measurements from the Ly- $\alpha$  forest (Wyithe & Bolton 2011; Calverley et al. 2011)
- \* Consistent with our definition of the “background onset time”

# The bubbles-in-island effect

(Xu et al. 2014)

\* Solving for a two-barrier problem:

1 - The first down-crossing distribution of random walks w.r.t. **island barrier**:

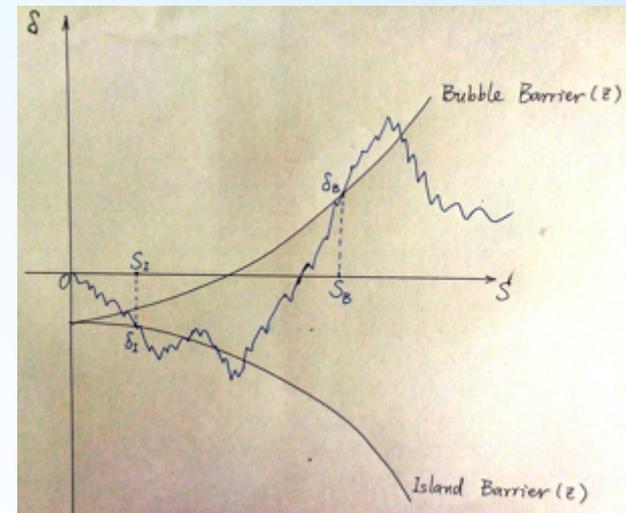
$$f_I(S_I, z)$$

2 - The conditional first up-crossing distribution w.r.t. **bubble barrier**:

$$f_B[S_B, \delta_B | S_I, \delta_I]$$

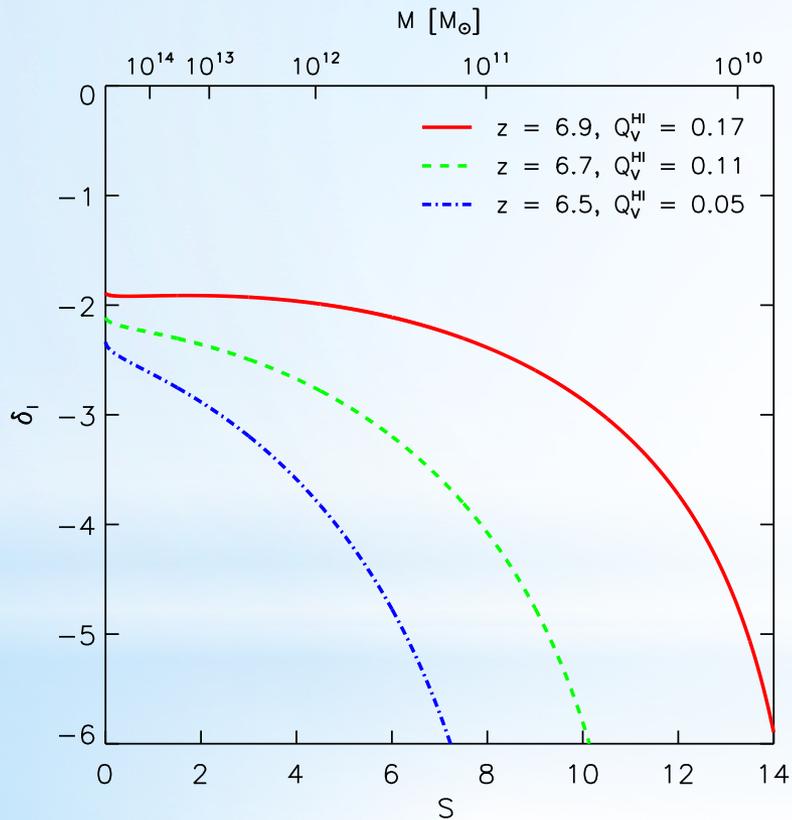
\* The effective bubble barrier:

$$\delta'_B = \delta_B(S + S_I) - \delta_I(S_I) \quad \text{where } S = S_B - S_I.$$

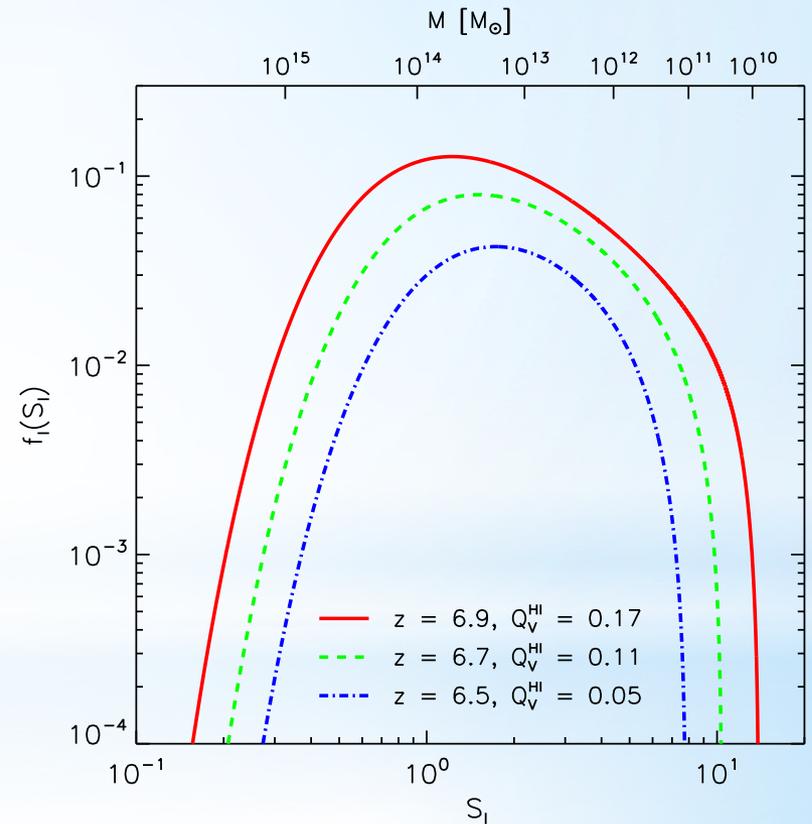


# The island-vS model - varying surface area

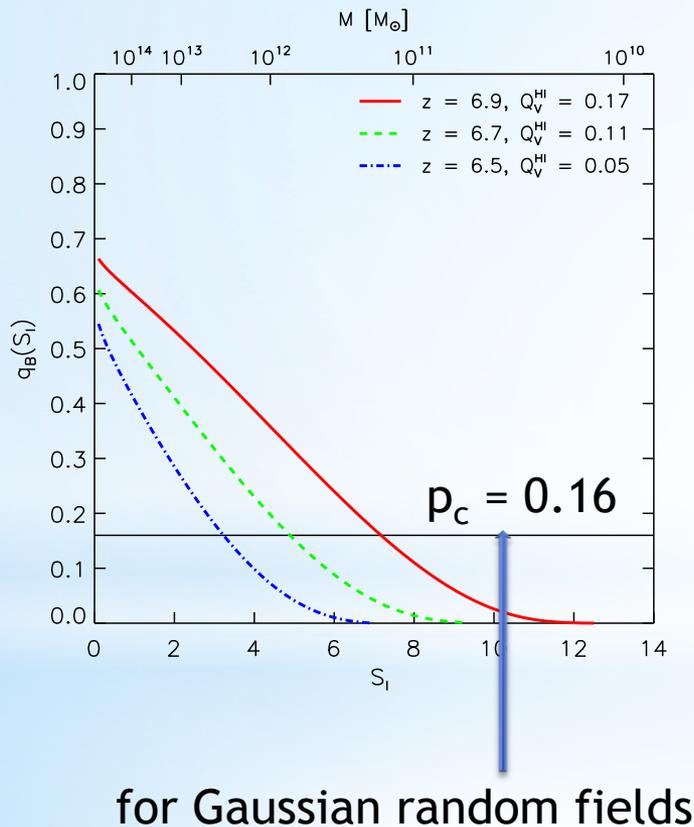
island barrier



first down-crossing distribution



# The problem of large bubbles-in-island fraction



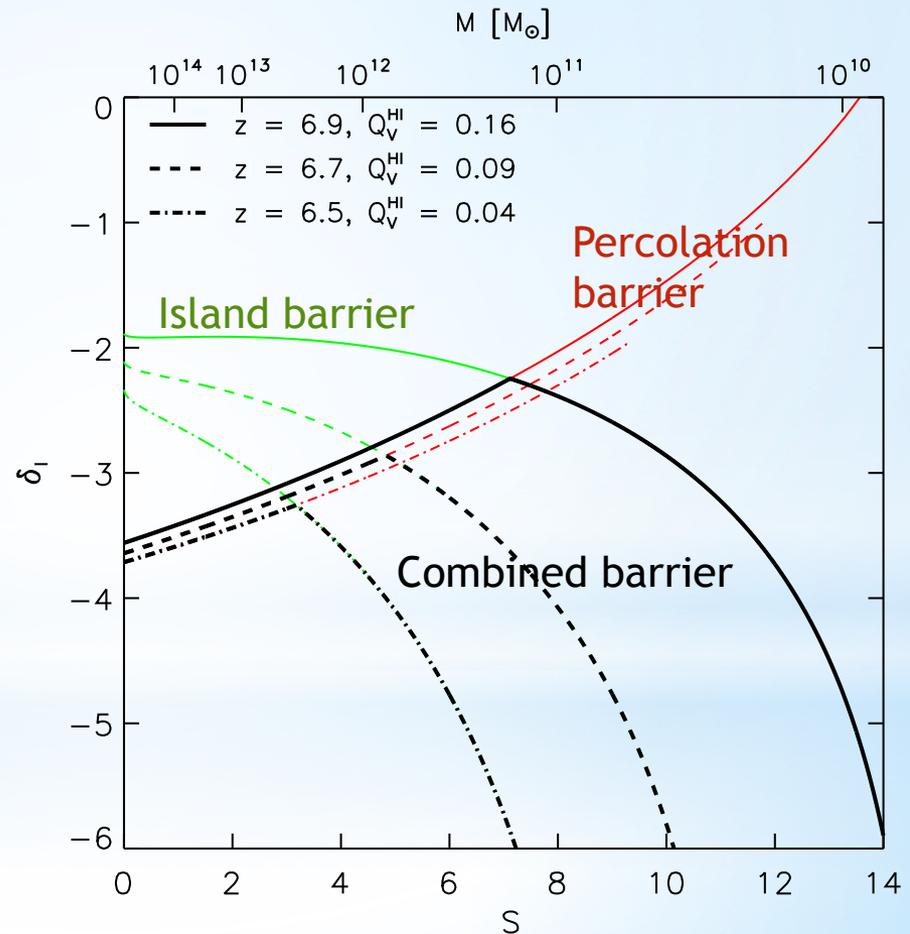
- \* Host islands  $\rightarrow$  overestimate the neutral fraction
- \* Neutral islands  $\rightarrow$  not the real image
- \* Difficult to visually identify the host islands
- \* Break down of bubble model inside islands

# The percolation criterion

\* The additional barrier is obtained by solving

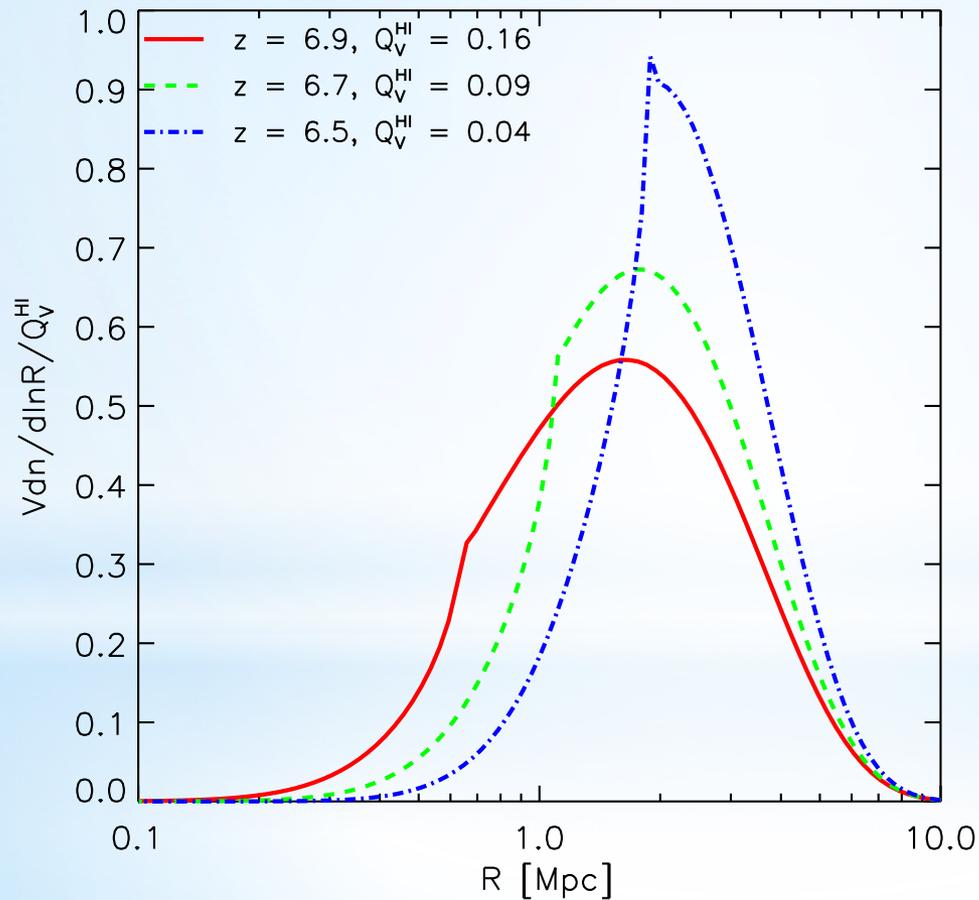
$$q_B(S_I, \delta_I; z) < p_c$$

\*  $p_c = 0.16$  for Gaussian random fields



# Results

- the size distribution with  $p_c$  cutoff



*A characteristic scale that does not change much with redshift!*

(Xu et al. 2014)

# Semi-numerical simulation - islandFAST

(Xu et al. 2017)

- \* Initial ionization field at  $z \gtrsim z_{\text{back}}$  generated by the 21cmFAST
- \* A *two-step* filtering algorithm
  - 1 - Based on the excursion set theory, we filter the evolved density field and *find host islands* with the island barrier including an ionizing background.
  - 2 - *Find bubbles in islands* with the bubble barrier without an ionizing background.
- \* A self-consistent treatment for the ionizing background taking into account the effect of absorption systems

$$\lambda_{\text{mfp}}^{-1}(z) = \lambda_{\text{I}}^{-1}(z) + \lambda_{\text{abs}}^{-1}(z).$$

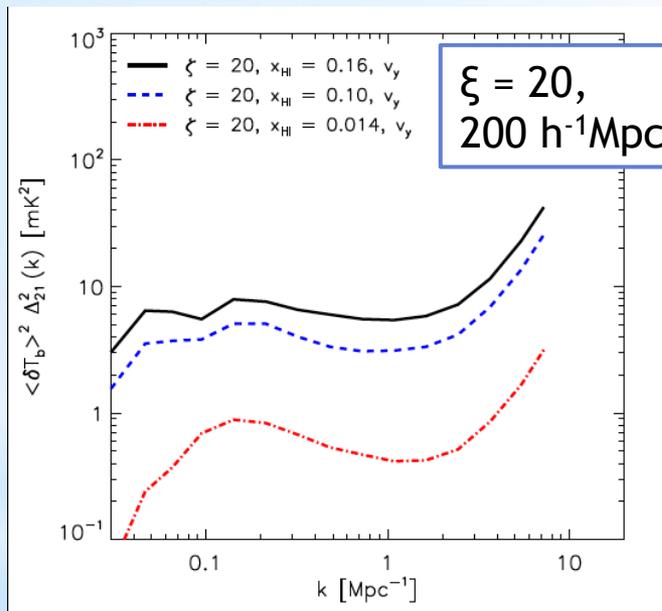
- \* An *iterative* procedure and *adaptive* redshift steps.

# Semi-numerical simulation - islandFAST

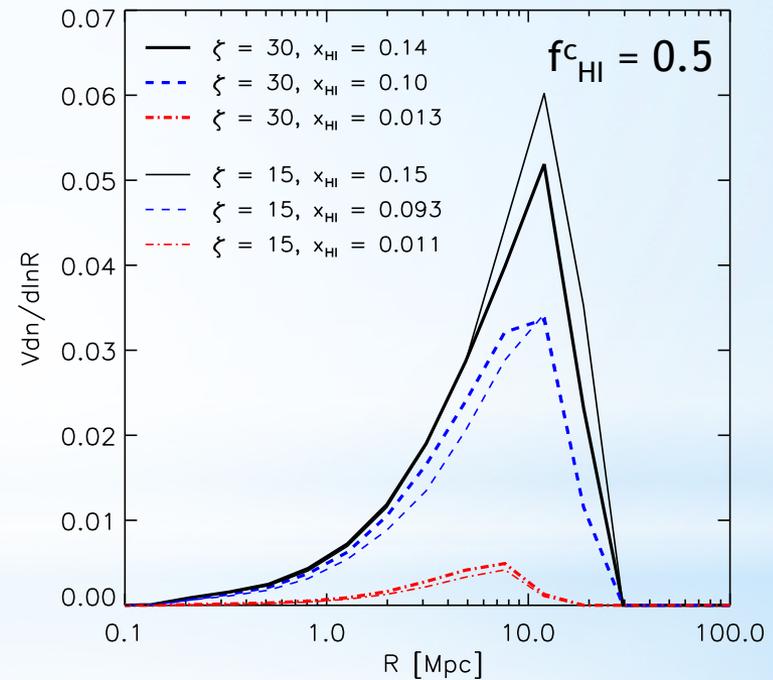


# Semi-numerical simulation - islandFAST

\* The 21 cm power spectrum

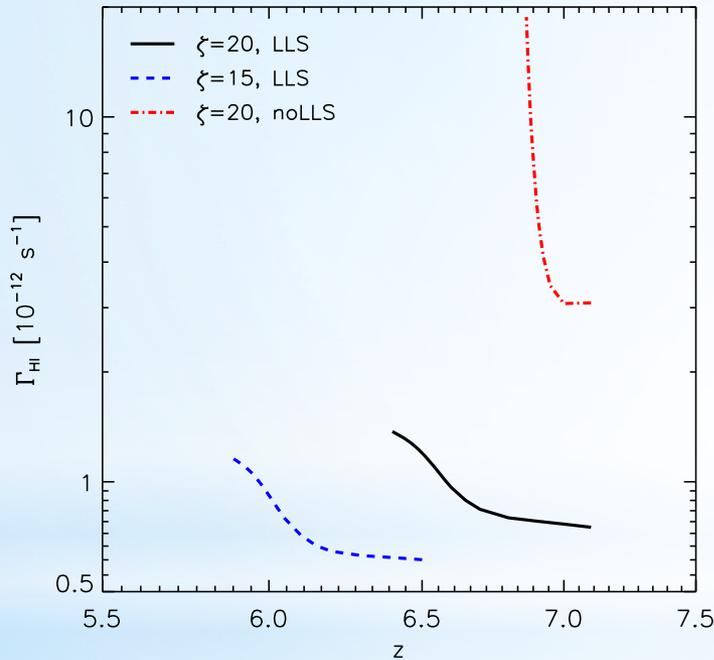


\* The size distribution of islands

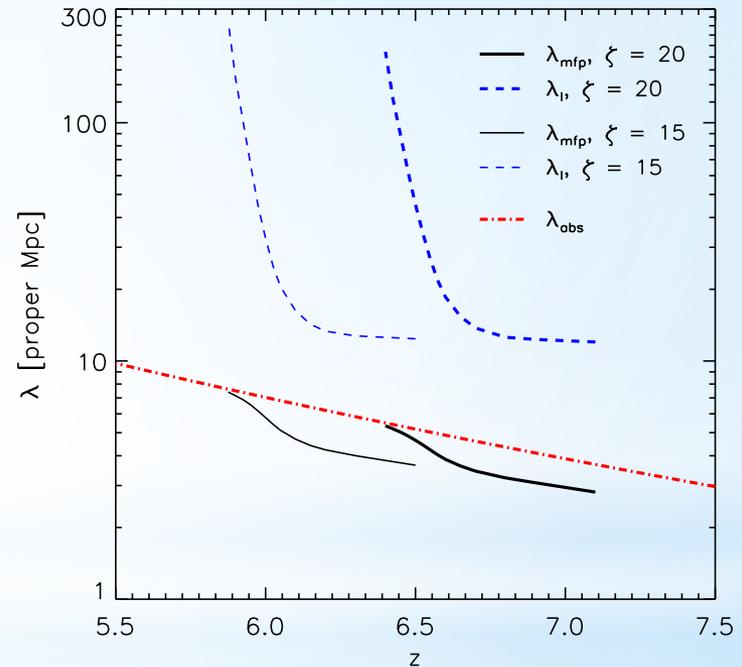


# The role of small absorbers vs. large islands

\* The ionizing background



\* The MFP of ionizing photons



- ✧ The large-scale islands dominate the morphology of the ionization field
- ✧ The small-scale absorbers dominate the opacity of the IGM, and they delay and prolong the reionization process significantly.

# Summary

## Before percolation

- \* Early EoR
- \* Isolated and spherical bubbles
- \* No UVB in model
- \* First-up-crossing distribution
- \* Linear-fitted barrier with analytical solution

## After percolation

- \* Late EoR
- \* Isolated and spherical islands
- \* With UVB
- \* First-down-crossing distribution
- \* Arbitrary shaped barriers with numerical solution
- \* Bubbles-in-island effect

\* **Bubble model vs. Island model**

# Summary

- islandFAST - a semi-numerical tool to ...
- \* Simulate the later part of EoR
- \* Model the effect of different absorbers
  - ✧ The large-scale islands dominate the morphology of the ionization field
  - ✧ The small-scale absorbers dominate the opacity of the IGM, and they delay and prolong the reionization process significantly.
- \* Generate synthesis signals from the late EoR
- \* .....

***THANK YOU!***