

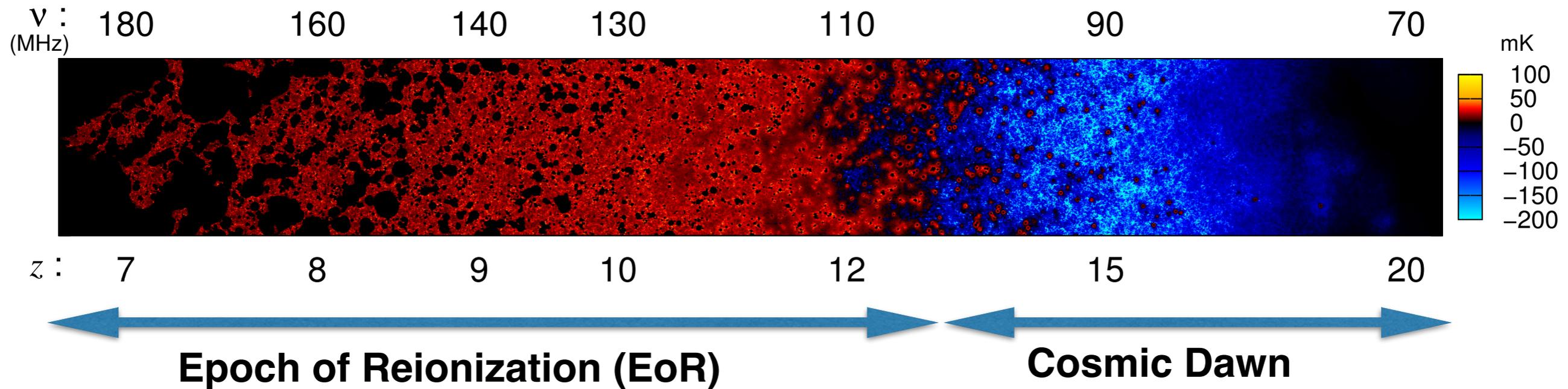
Towards simulating and quantifying the light-cone EoR/CD 21-cm signal

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Collaborators

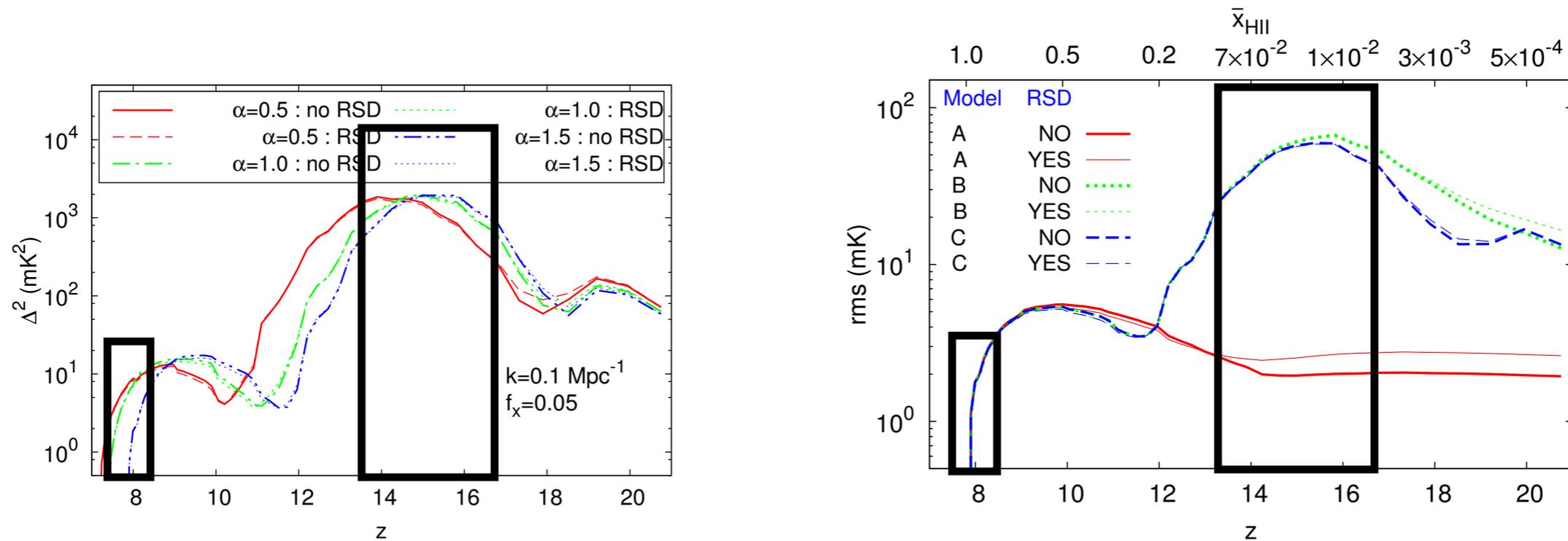
- Garrelt Mellema
- Hannes Jensen
- Suman Majumdar
- Ilian T. Iliev
- Raghunath Ghara
- Tirthankar Roy Choudhury
- Rajesh Mondal
- Somnath Bharadwaj

Light cone effect on 21-cm signal from EoR and Cosmic Dawn



- 3D imaging of universe with 21-cm observations
- The mean and statistical properties of HI 21-cm signal change with redshift.
- This effect, known as the ‘light-cone’ (LC) effect.
- LC effect has a significant impact on 3D power spectrum etc.

Light cone effect on 21-cm signal: An example



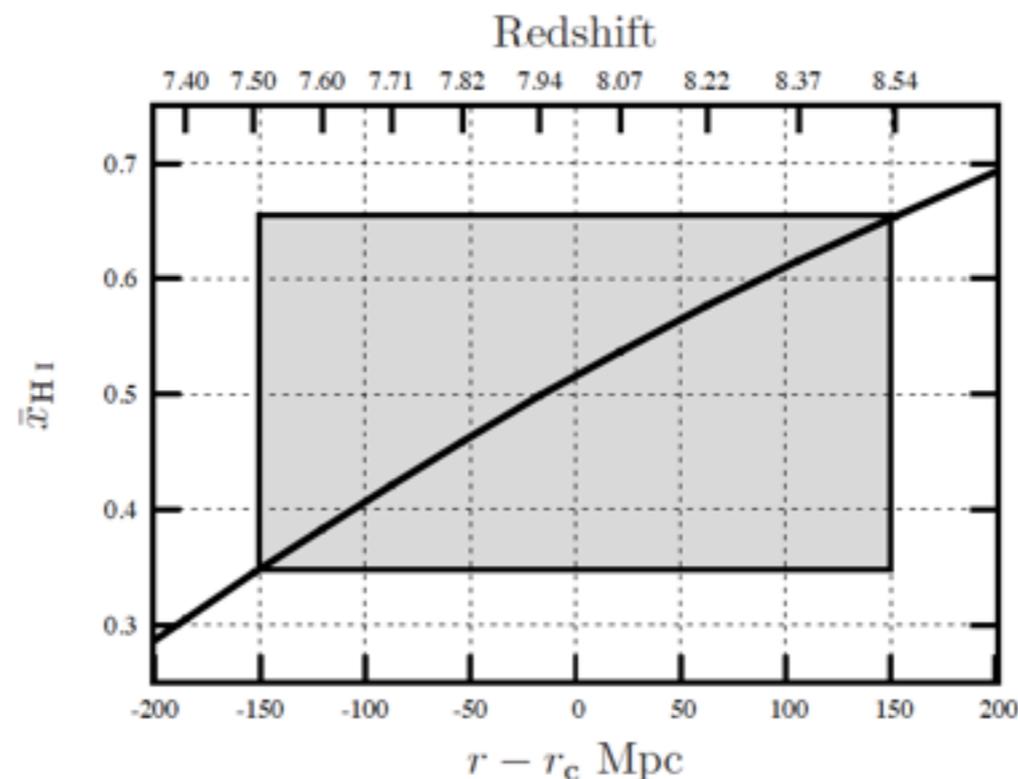
16 MHz frequency bandwidth covers redshift range
 $\sim 7.5 < z < 8.5$ (at $z=8$)
 $\sim 13.4 < z < 16.6$ (at $z=15$)

Two important issues

- How to simulate Light Cone 21-cm signal?
- How to quantify the Light Cone signal ?

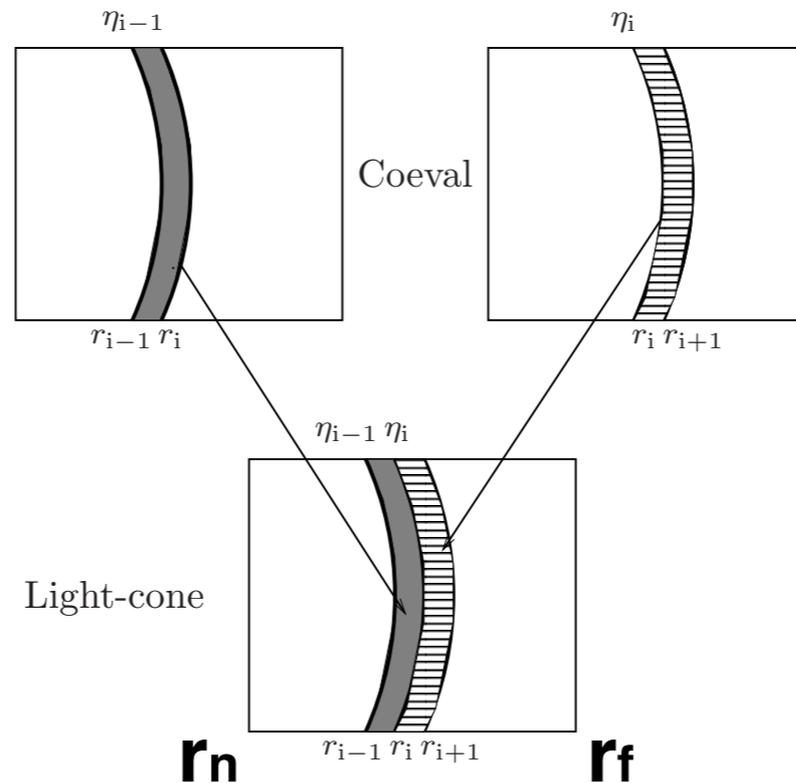
Generating the coeval cubes

- Used semi-numerical simulations (Choudhury et al (2009), Majumdar et al(2012), Mondal et al (2016))
- Box size $L = 300.16$ Mpc
- 4288^3 particles, grids of spacing 0.07 Mpc
- Mass resolution of $\sim 10^8 M_{\odot}$
- Identify collapsed halos using FoF
- generated the coeval cubes in the range $r_n = 9001.45$ Mpc (nearest) to $r_f = 9301.61$ Mpc (farthest) ($7.51 < z < 8.53$).
- The change in the mass-averaged HI fraction \bar{x}_{HI} is $\bar{x}_{\text{HI}} \approx 0.65 - 0.35 = 0.30$
- 25 different coeval simulations.
- The LC box is centred at redshift 8
- co-moving distance $r_c = 9151.53$ Mpc, frequency $\nu_c = 157.78$ MHz and $\bar{x}_{\text{HI}} \approx 0.51$.



Simulating Light Cone 21-cm signal -A particle based method

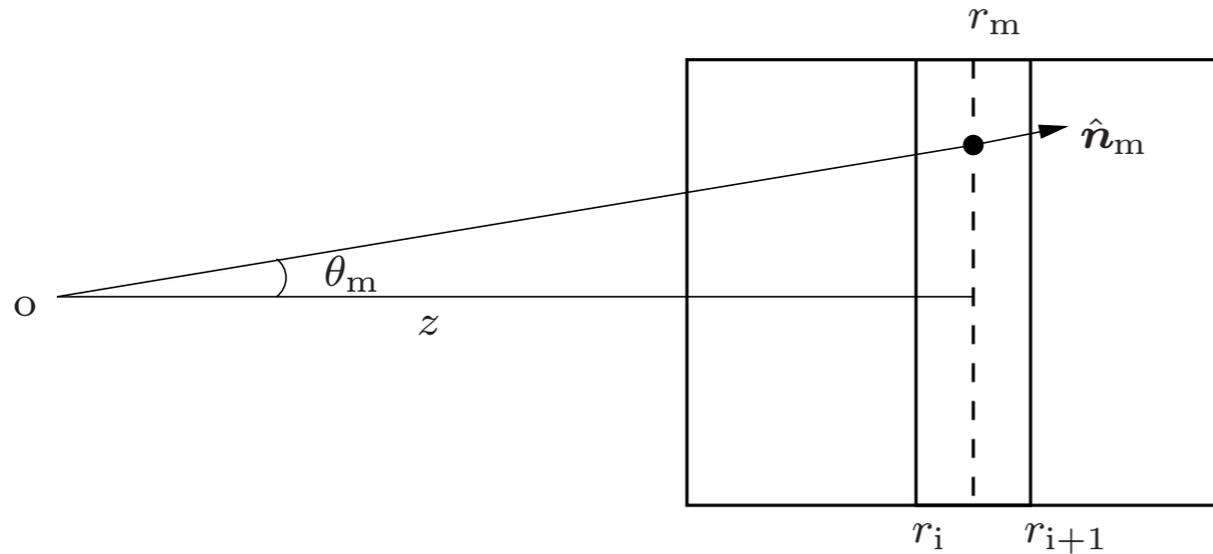
- We consider a simulation that span the comoving distance range \mathbf{r}_n (nearest) to \mathbf{r}_f (farthest).



- The HI distribution is represented by particles
- HI masses vary with position
- For each slice we have filled the region r_i to r_{i+1} with the HI particles from the corresponding region in the coeval box

Flat sky approximation

Mapping to angle (theta) -frequency coordinates



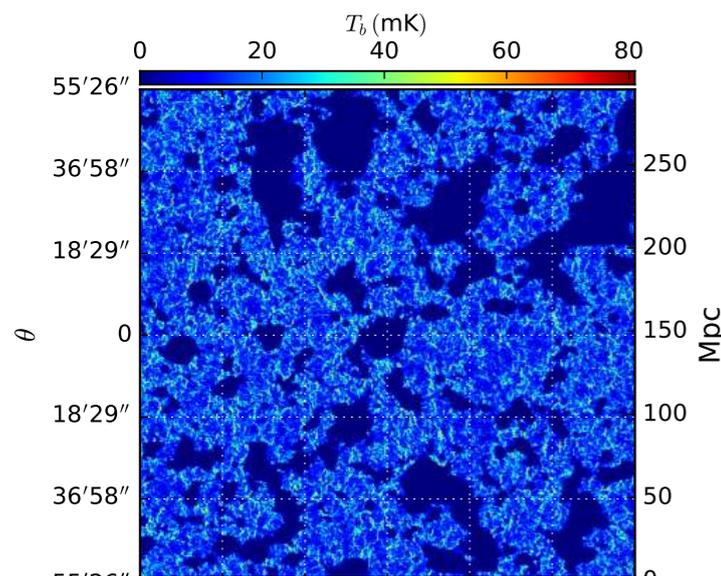
$$\hat{n} = \hat{k} + \theta$$

$$\theta_x = x/r_c, \theta_y = y/r_c \quad \text{eq. 1}$$

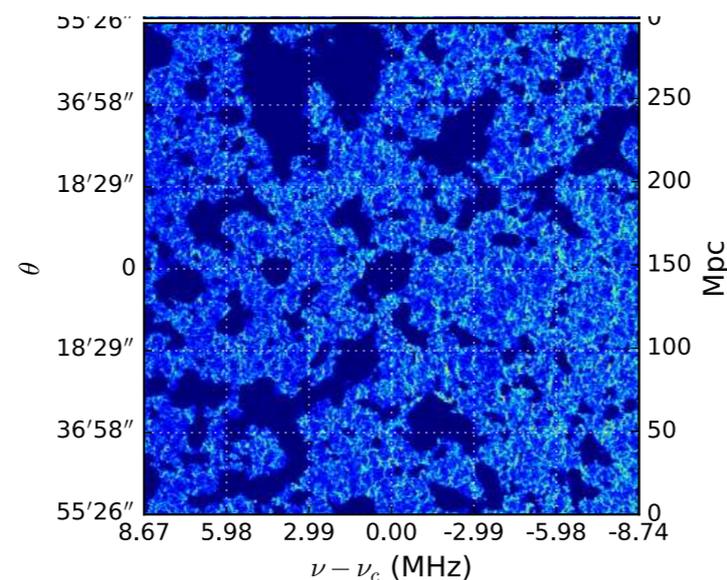
$$\nu_m = \nu_i \left[1 - \frac{a_i H_i (r_m - r_i) + \hat{n}_m \cdot \mathbf{v}_m}{c} \right] \quad \text{eq. 2}$$

$$\rho'_{\text{HI}} = (\Delta\Omega \Delta\nu)^{-1} \left(\frac{H_0 \nu_c}{c} \right) \sum_m \frac{[M_{\text{HI}}]_m}{r_n^2} \quad \text{eq. 3}$$

$$T_b(\hat{n}, \nu) = \bar{T}_0 \frac{\rho'_{\text{HI}}}{\bar{\rho}_{\text{H}}} \quad \text{eq. 4}$$



Coeval



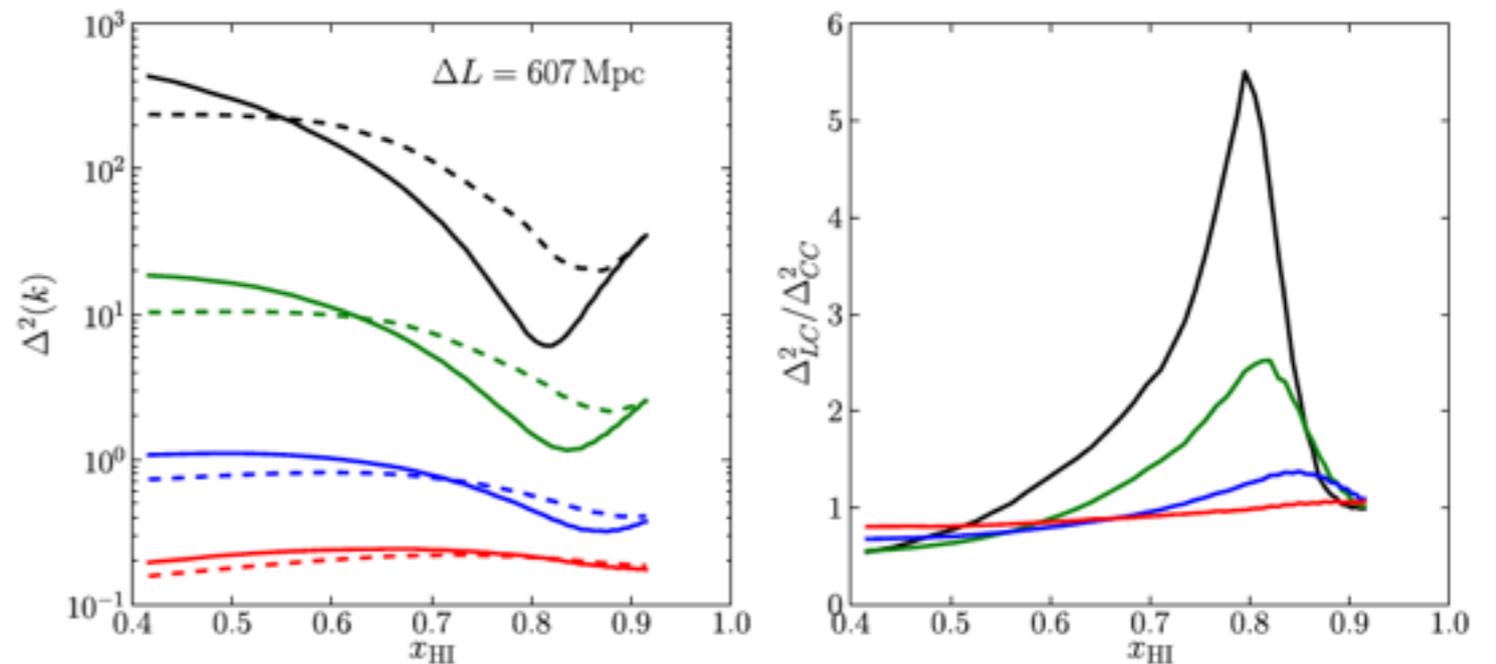
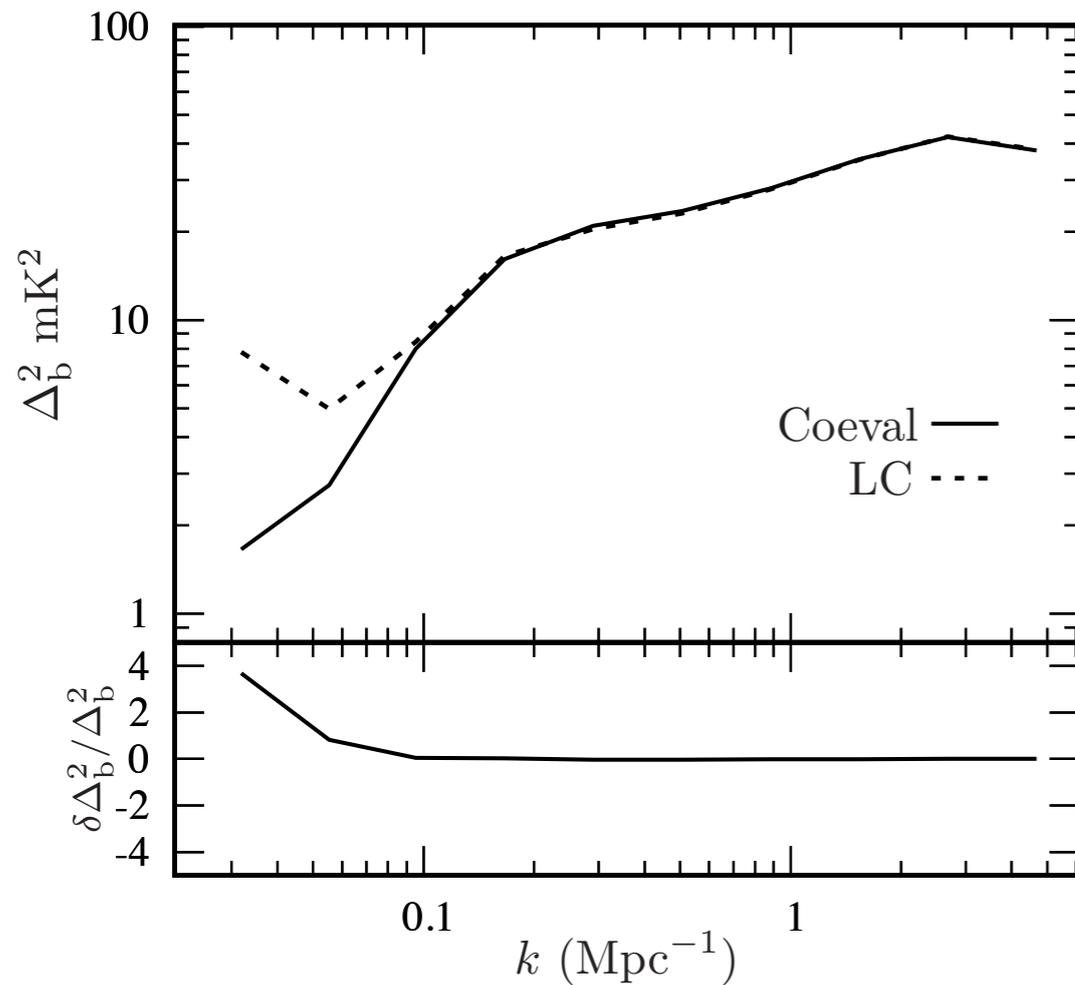
Light cone

LC effect has been implemented before the peculiar velocity effect- **very important!!!**

Effects on EoR 3D power spectrum

Reverse mapping of EoR 21-cm brightness temperature fluctuations to spatial comoving coordinates $\delta T_b(\theta_x, \theta_y, \nu) \rightarrow \delta T_b(x, y, z)$

$$(x, y, z) = [r_c \theta_x, r_c \theta_y, z_c + r_c' (\nu - \nu_c)]$$

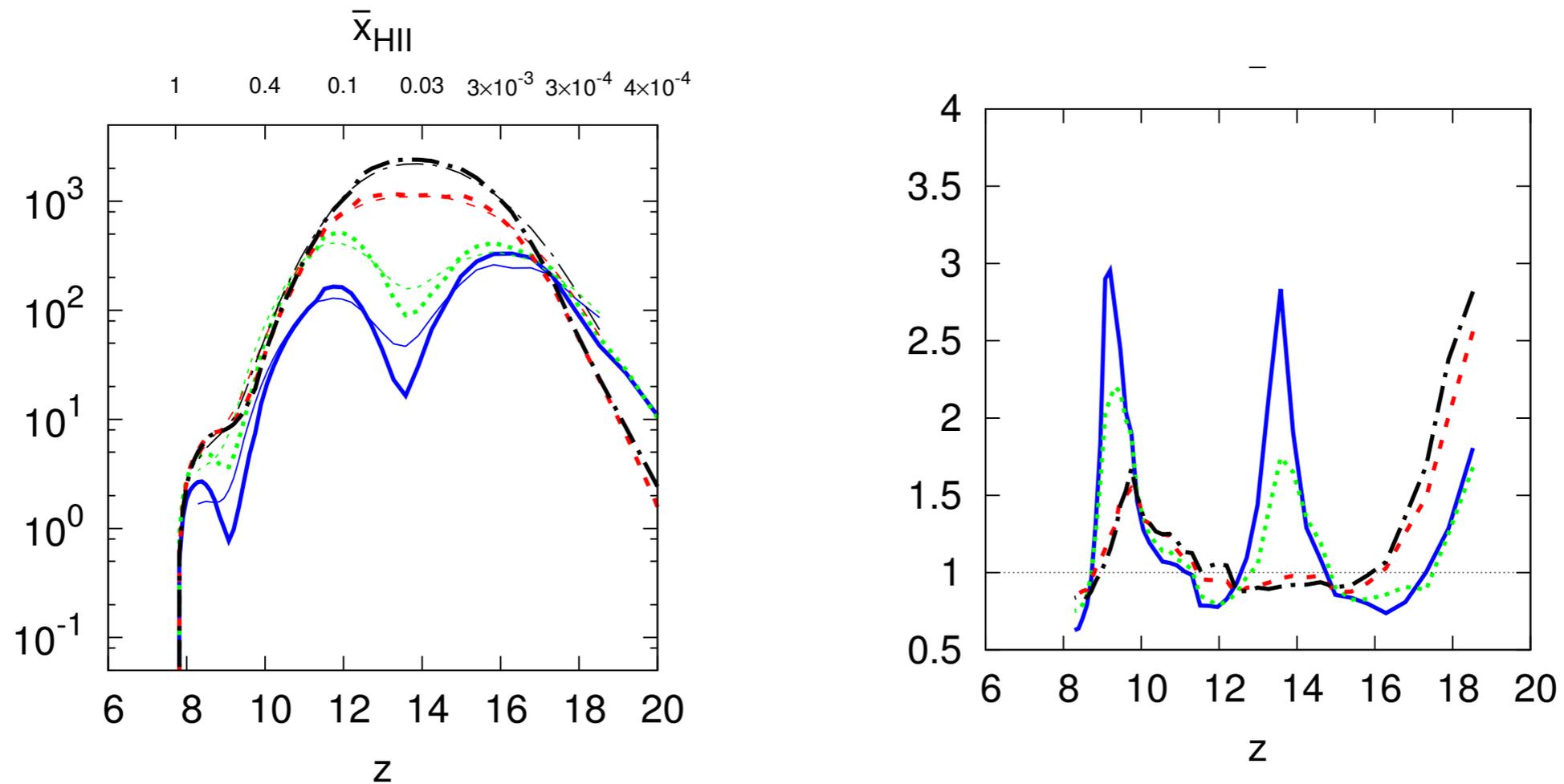


The LC effect is stronger at large scales.

Mondal, Bharadwaj, Datta, 2017

Datta, Jensen, Mellema et al 2014

Effects on Cosmic dawn 21-cm power spectrum



- The LC effect becomes when power spectrum takes a turn.
- Peaks and dips are smoothed out due to the LC effect

Ghara, Datta, Choudhury, 2015

But

Can the 3D power spectrum describe light-cone 21-cm signal completely?

Mondal, Bharadwaj, Datta, 2017

Light Cone effect breaks the statistical homogeneity

and

makes the signal non-ergodic

along the LoS.

In addition

Light Cone effect breaks periodicity of the signal along LOS

Information theory

- The Fourier modes form orthogonal Eigen basis only for a statistically homogeneous or ergodic signal
- The second order statistics is completely quantified by the 3D power spectrum $P(k)$.
- 3D $P(k)$ assumes that the signal is ergodic (E) and periodic (P) along the LoS direction. Both are not valid for Light cone 21-cm signal
- For light cone signal which is non-ergodic, **different Fourier modes along the LoS is correlated** and **$P(k)$ does not retain the entire information of the 21-cm signal.**

The multi-frequency angular power spectrum (MAPS) $\mathcal{C}_\ell(\nu_1, \nu_2)$

Datta, Choudhury, Bharadwaj, 2007

$$\delta T_b(\hat{\mathbf{n}}, \nu) = \sum_{\ell, m} a_{\ell m}(\nu) Y_\ell^m(\hat{\mathbf{n}})$$

$$\mathcal{C}_\ell(\nu_1, \nu_2) = \langle a_{\ell m}(\nu_1) a_{\ell m}^*(\nu_2) \rangle$$

In the flat-sky approximation

$$\mathcal{C}_\ell(\nu_1, \nu_2) = \mathcal{C}_{2\pi U}(\nu_1, \nu_2) = \Omega^{-1} \langle \tilde{T}_{b2}(\mathbf{U}, \nu_1) \tilde{T}_{b2}(-\mathbf{U}, \nu_2) \rangle$$

$\tilde{T}_{b2}(\mathbf{U}, \nu)$ is the 2D Fourier transform of $\delta T_b(\boldsymbol{\theta}, \nu)$

By imposing the conditions (i) $\mathcal{C}_\ell^{\text{EP}}(\nu_1, \nu_2) = \mathcal{C}_\ell^{\text{EP}}(\Delta\nu)$ (ergodic)

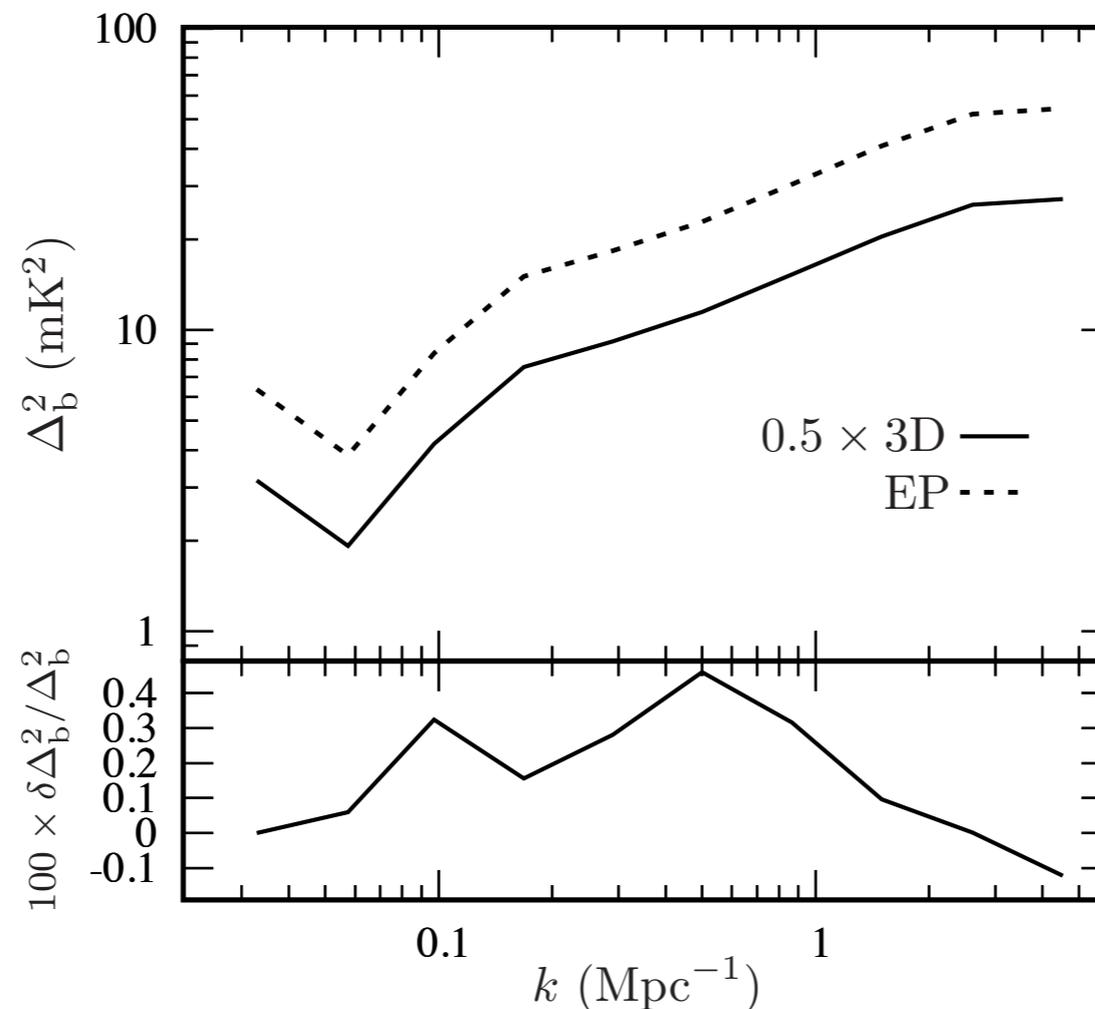
(ii) $\mathcal{C}_\ell^{\text{EP}}(\Delta\nu) = \mathcal{C}_\ell^{\text{EP}}(B - \Delta\nu)$ (periodic)

We show

$$P(k_\perp, k_\parallel) = r_c^2 r'_c \int d(\Delta\nu) e^{-ik_\parallel r'_c \Delta\nu} \mathcal{C}_\ell^{\text{EP}}(\Delta\nu)$$

If ergodicity and periodicity are imposed, 3D power spectrum is exactly recovered from MAPS.

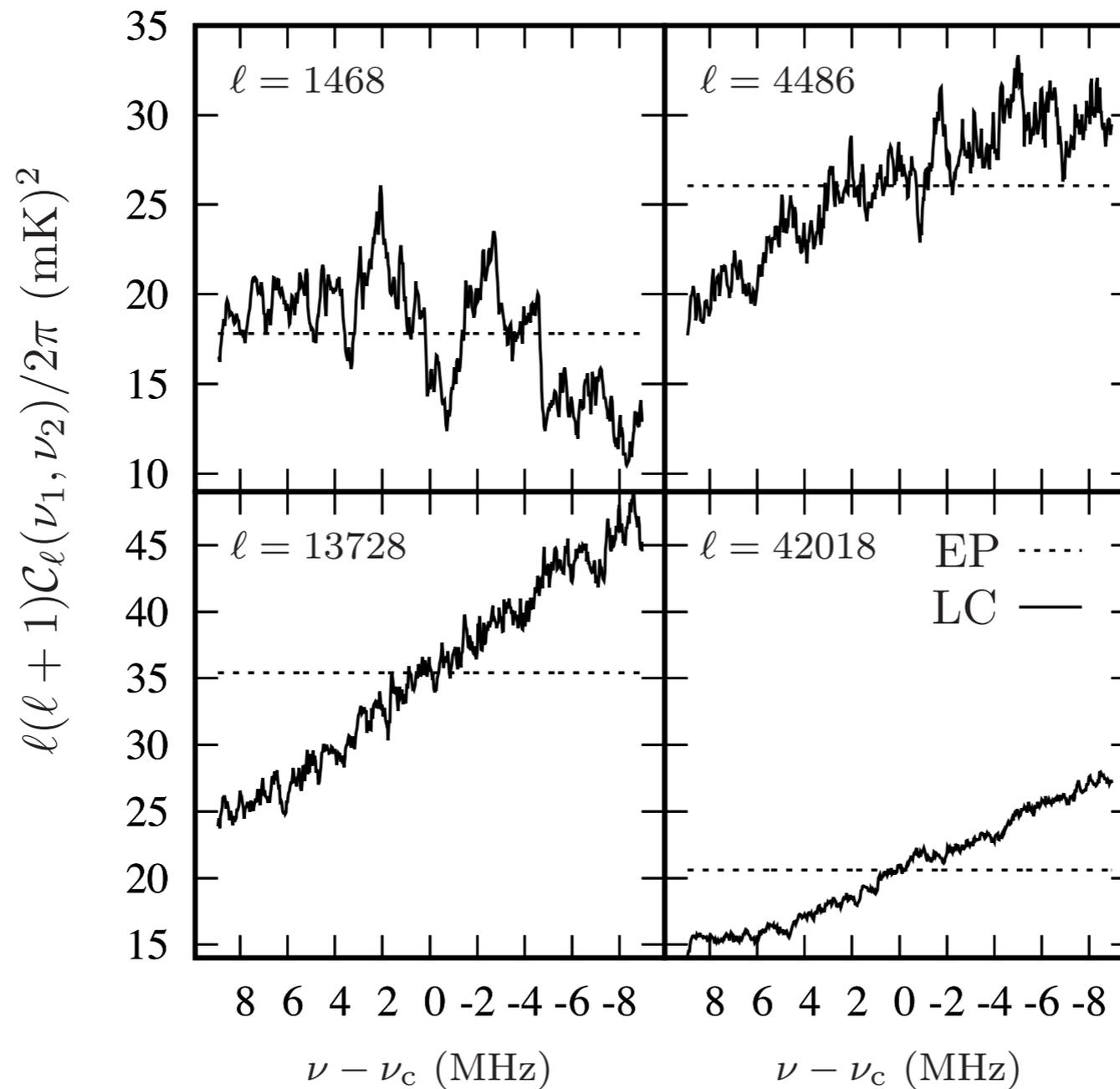
$$P(k_{\perp}, k_{\parallel}) = r_c^2 r'_c \int d(\Delta\nu) e^{-ik_{\parallel} r'_c \Delta\nu} \mathcal{C}_{\ell}^{\text{EP}}(\Delta\nu)$$



This implies that MAPS contains much more information than $P(k)$ and a better statistics to quantify light cone EoR 21-cm signal.

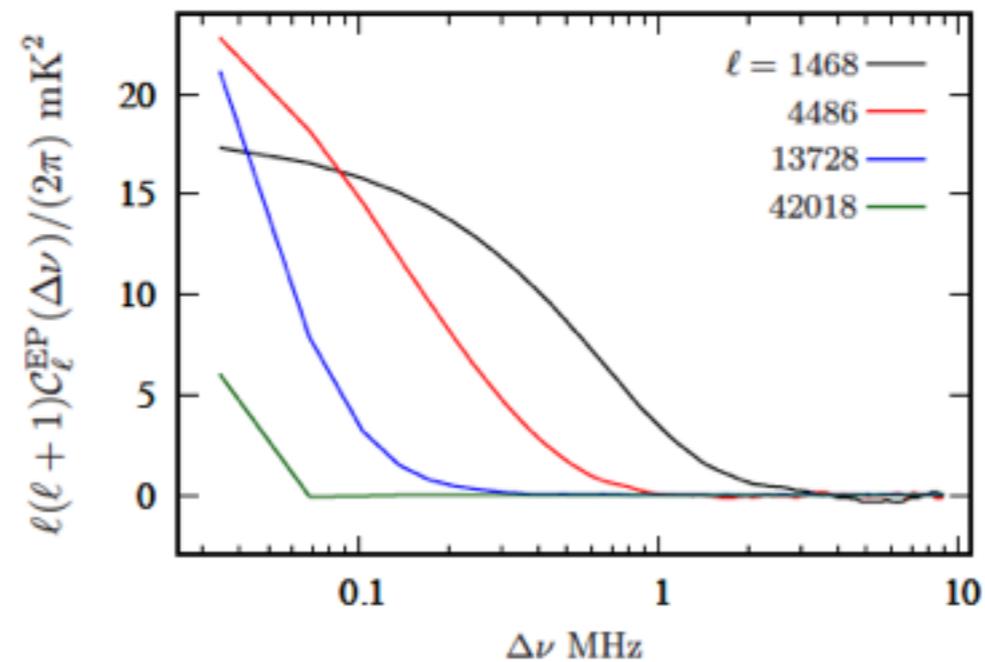
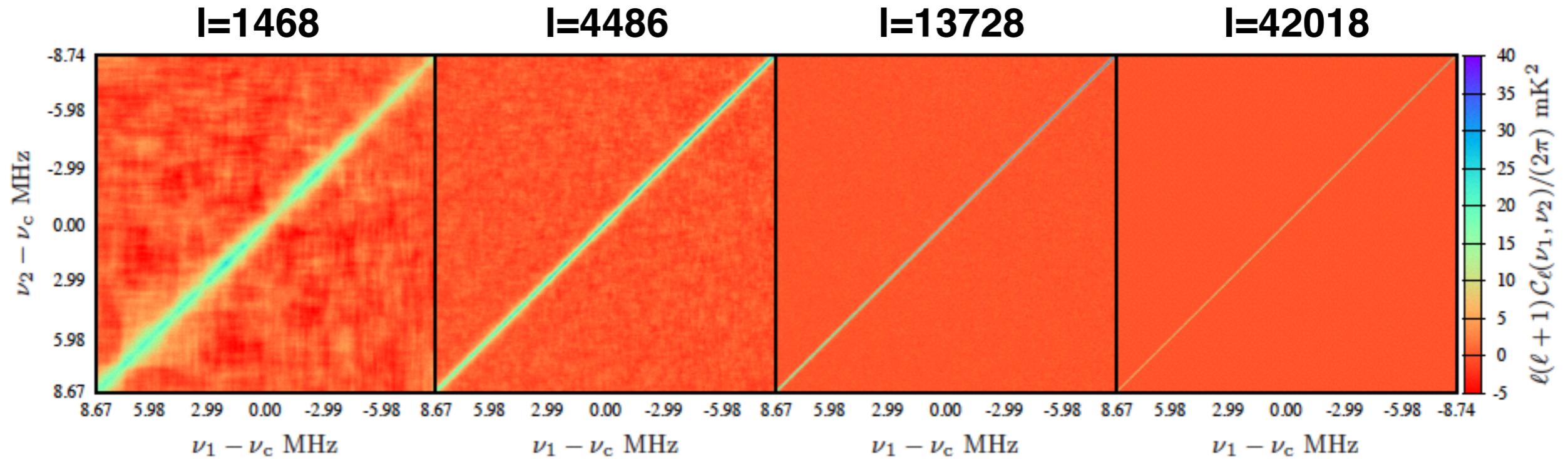
Results -MAPS

From Light cone and coeval simulations



Results -MAPS

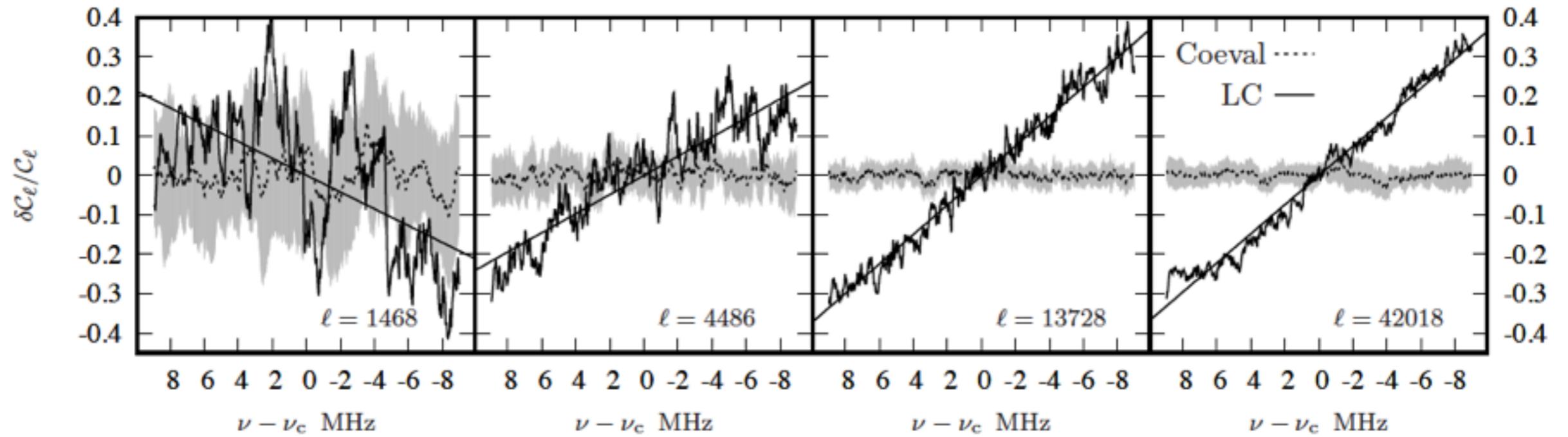
From Light cone simulations



Assuming periodicity
and ergodicity

Results -MAPS

From Light cone and coeval simulations



Summary

- The light cone effect is found to be an important effect and should be considered in the statistical analysis of EoR 21-cm signal.
- It is important that the LC effect should be implemented before correcting for the peculiar velocity effect.
- The LC effect makes the signal non-ergodic and non-periodic along the los.
- 3D power spectrum or spherically averaged power spectrum of HI 21-cm signal can not retain full information of the light cone EoR/CD signal.
- Multi-frequency angular power spectrum (MAPS) retains the entire information of the second order statistics of the signal and therefore a better statistical estimator of the signal. The 3D power spectrum can be recovered from the MAPS assuming the signal to be ergodic and periodic along the los.

Thank you